

1 **Entropy Production and Coarse Graining of the Climate Fields in**  
2 **a General Circulation Model**

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## Abstract

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10 We extend the analysis of the thermodynamics of the climate system by investigating the role  
11 played by processes taking place at various spatial and temporal scales through a procedure of  
12 coarse graining. We show that the coarser is the graining of the climatic fields, the lower is  
13 the resulting estimate of the material entropy production. In other terms, all the spatial and  
14 temporal scales of variability of the thermodynamic fields provide a positive contribution to the  
15 material entropy production. This may be interpreted also as that, at all scales, the temperature  
16 fields and the heating fields resulting from the convergence of turbulent fluxes have a negative  
17 correlation, while the opposite holds between the temperature fields and the radiative heating  
18 fields. Moreover, we obtain that the latter correlations are stronger, which confirms that radiation  
19 acts as primary driver for the climatic processes, while the material fluxes dampen the resulting  
20 fluctuations through dissipative processes. We also show, using specific coarse-graining procedures,  
21 how one can separate the various contributions to the material entropy production coming from  
22 the dissipation of kinetic energy, the vertical sensible and latent heat fluxes, and the large scale  
23 horizontal fluxes, without resorting to the full three-dimensional time dependent fields. We find that  
24 most of the entropy production is associated to irreversible exchanges occurring along the vertical  
25 direction, and that neglecting the horizontal and time variability of the fields has a relatively small  
26 impact on the estimate of the material entropy production. The approach presented here seems  
27 promising for testing climate models, for assessing the impact of changing their parametrizations  
28 and their resolution, as well as for investigating the atmosphere of exoplanets, because it allows for  
29 evaluating the error in the estimate of their thermodynamical properties due to the lack of high-  
30 resolution data. The findings on the impact of coarse graining on the thermodynamic fields on  
31 the estimate of the material entropy production deserve to be explored in a more general context,  
32 because they provide a way for understanding the relationship between forced fluctuations and  
33 dissipative processes in continuum systems.

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## 34 I. INTRODUCTION

35 Along the lines of the theoretical construction due to Lorenz [25, 26] of energy cycle of the  
36 atmosphere, the climate can be seen as a non-equilibrium multi-scale system, which generates  
37 entropy through a variety of irreversible processes [8, 30, 39–41], and transforms moist  
38 static energy into mechanical energy, as it features a positive spatio-temporal correlation  
39 between heating and temperature patterns, so that it can be represented schematically as a  
40 heat engine with a given efficiency [14, 28]. For a given value of the external and internal  
41 parameters, the climate system achieves a steady state by balancing the input and output  
42 of energy and entropy with the surrounding environment [42]. The large scale motions of  
43 the geophysical fluids are at the same time the result of the mechanical work produced by  
44 the climatic engine, and contribute to reducing the temperature gradients which make the  
45 energy conversion possible [41, 48]. Obtaining a closure to this problem would be equivalent  
46 to developing a self-consistent theory of climate dynamics.

47 Developing a comprehensive theory of climate dynamics is one of the grand contemporary  
48 scientific challenges, also for its obvious environmental, social and economical relevance,  
49 and it is far from being an accomplished task [11, 45]. In recent years, the extraordinary  
50 developments of planetary sciences coming from the discovery of extra-solar planets and the  
51 ensuing need for understanding the properties of atmospheric circulations realized under  
52 physical and chemical conditions very different from those of the Earth and of the other  
53 solar planets has provided further stimulation in this direction [46].

54 While the thermodynamic interpretation of the baroclinic disturbances, which provide  
55 the dominant contributions to the low-to-high latitudes heat transport, lies at the core of  
56 dynamical meteorology [13], and thermodynamics provides indeed the best framework for  
57 studying strong meteorological features like hurricanes [6], recent results suggest that the  
58 structural properties of the climate system [44] and in particular its tipping points [23, 35]  
59 can be effectively analyzed using the thermodynamic indicators developed in [28], with the  
60 efficiency and the entropy production providing the most interesting indicators [1, 31, 32].  
61 Moreover, recent studies have underlined that it is instead possible to define generalized  
62 climate sensitivities able to describe quite accurately the responses of thermodynamic quan-  
63 tities to changes in CO<sub>2</sub> concentration [29].

64 Despite its relevance at theoretical level [2, 20], traditionally, entropy production is not

65 one of the first physical quantities climate modelers investigate when assessing the perfor-  
66 mance of a global climate model (GCM) or the response of the climate system to forcings.  
67 One should note that the attention towards entropy production in the climate system and  
68 in climate modeling has been revived when several authors started proposing it as target  
69 function to maximize when tuning free or empirical parameters of approximate numerical  
70 models [17, 22, 27] or for getting good first order approximations of the climate state without  
71 resorting to long integrations [12, 35] This is the *weak* or *pragmatic* version of the so-called  
72 maximum entropy production principle (MEPP) [18], which, in its *strong* form, proposes  
73 that any non-equilibrium systems adjust itself in order to maximize the entropy production;  
74 see [18, 34]. The MEPP theoretical foundations [3, 4] have been criticized both at theoretical  
75 level [10] and in terms of its geophysical applications [9, 37], so that weaker formulations  
76 are now mostly preferred [5]. Recently, some authors have turned their attention on testing  
77 whether it is possible to propose a a variational principle for applies to another index of the  
78 irreversibility of the system, namely the rate of dissipation of kinetic energy [19, 38].

79 In this paper, we wish to investigate the entropy production of a climate model for  
80 studying, instead of large scale balances, its fluctuations at different temporal and spatial  
81 scales. Climate is a multi-scale system where dynamics takes place on vast range of inter-  
82 acting scales. The definition of parametrizations for unresolved scales is a major challenge  
83 of climate modeling and the proposal of closure theories connecting small and large scale  
84 properties is a major part of any attempt at formulating approximate theories for climate  
85 dynamics. The issue of understanding the impact of small scales on large scales and vice-  
86 versa, and of performing properly the upscaling and downscaling of a model's output is  
87 of great relevance also for intercomparing the performances of various versions of a given  
88 numerical model, or of a set of numerical model simulating the same system, differing for  
89 the adopted spatial and temporal resolution, and for comparing model data to observations.

90 Our goal is manifold. One one side, we want to introduce a way to evaluate how the  
91 different scales of motion contribute to the overall entropy production of the climate system.  
92 This investigation, therefore, complements the investigation of how much energy is contained  
93 in the various scales of motion and of the energy fluxes across these scales. In order to achieve  
94 this goal, we consider the entropy budget of the FAMOUS GCM [15] in standard, present  
95 climate configuration, taking advantage of the fact that it is one of the very few climate  
96 models where the entropy production diagnostics has been implemented and throughly tested

97 [36, 37]. Starting from the output fields at the highest possible resolution given by the  
98 model (temporal resolution of one time step, and same spatial resolution of the actual  
99 numerical model), we perform a coarse graining in space and in time to the dynamical and  
100 thermodynamical fields appearing in the terms describing the entropy production of the  
101 system, and we test how the estimate of the entropy production changes when different  
102 coarse graining are applied to the data. We anticipate that we obtain that the coarser is  
103 the graining procedure, the lower is the estimate of the entropy production one obtains, as  
104 somehow intuitive. One must note that this is, in fact an obvious result when considering  
105 simple diffusive system, but not so obvious when fully nonlinear, multiphase systems are  
106 considered. We also obtain similar results when considering different degrees of longitudinal  
107 averaging of the fields, up to considering zonally averaged fields only. Our findings provide  
108 a way to assess how having low-resolution information about the dynamics of turbulent  
109 systems affects our ability to reconstruct its thermodynamical properties. Moreover, the  
110 procedure discussed in this paper allows to put on firmer ground the results proposed in  
111 [30] on the possibility of separating vertical and horizontal exchange processes as far as  
112 entropy production is concerned. Finally, we can study in detail the relationship between  
113 two apparently equivalent ways of computing the entropy production proposed in [8].

114 This paper is structured as follows. In section II we briefly recapitulate some definitions  
115 and equations relevant for setting the problem of computing the entropy production of the  
116 climate system we explain how to perform such a calculation in a climate model. In section  
117 III we explain what we mean precisely by coarse graining of the data and describe how it  
118 is actually implemented in the model's output. We also provide some conjectures what will  
119 be discussed in later in the paper. In section IV we present our results. We first describe  
120 the impact of performing coarse graining on time alone, thus exploring the range between  
121 time step data and long term averaged data, and then we extend our analysis to the space  
122 domain, showing how performing zonal, horizontal, and mass-weighted averaging over the  
123 output data impacts the obtained estimate of the entropy production. In section V we  
124 present our conclusions and perspective for future works. In appendix A we present some  
125 theoretical arguments on a simple diffusive system for clarifying the meaning of the results  
126 obtained from the data analysis.

## 127 II. CLIMATE ENTROPY BUDGET

128 Following [2, 20], for any system it is possible to decompose the rate of change of its  
129 entropy  $dS/dt$  as  $dS/dt = d_e S/dt + d_i S/dt$ , where the first term is called the external and  
130 the second term is the internal contribution to the entropy budget. The external contribution  
131 corresponds to the entropy flux through the boundaries of the system whereas the internal  
132 entropy production is associated with the irreversible processes taking place in the system.  
133 The second law of thermodynamics imposes that the internal entropy production has to  
134 be nonnegative at all instants, so that  $d_i S/dt \geq 0$ . When a statistically steady state is  
135 achieved, the external and internal entropy production have to balance each other so that  
136 the total rate of entropy change is zero: we have  $\overline{dS/dt} = 0 \rightarrow \overline{d_i S/dt} = -\overline{d_e S/dt} \geq 0$ , where  
137 the overline indicates averaging over a long time interval compared to the internal scales of  
138 the system. The previous expression means that a non-equilibrium system generates on  
139 the average a positive amount of entropy through irreversible processes, and such excess  
140 of entropy is expelled at the boundaries. Non-equilibrium is maintained if the system is  
141 in contact with more than one reservoir with given temperature and/or chemical potential  
142 [7]. Of course, if the system is at equilibrium, the previous inequality becomes an equality,  
143 as in the long run the system reaches an homogeneous state of maximum entropy and no  
144 additional entropy is generated.

145 In the climate system two rather different set of processes contribute to the total entropy  
146 production [8, 41]. The first set of processes are responsible for the irreversible thermaliza-  
147 tion of the photons emitted near the Sun's corona at roughly 5700 K at the much lower  
148 temperatures, typical of the Earth's climate. This contributes for about 95% of the total  
149 average rate of entropy production for our planet, which is about  $0.90 \text{ W m}^{-2}$  [8, 41]. The  
150 remaining contribution is due to the processes responsible for mixing and diffusion inside the  
151 fluid component of the Earth system, and for the dissipation of kinetic energy due to viscous  
152 processes. This constitutes the so-called material entropy production, and is considered to  
153 be the entropy related quantity of main interest as far as the properties of the climate system  
154 are concerned. See [37] for an extensive discussion of this issue a careful estimate of its value  
155 in two climate models, including the one used in this study.

156 When separating the entropy budget for radiation and for the fluid part of the climate

157 system, and taking long term averages, one can derive the following equation [8, 14]:

$$\int_V d^3\mathbf{x} \left[ \overline{\left(\frac{\dot{q}_{rad}}{T}\right)} + \overline{\dot{s}_{mat}} \right] = 0 \quad (1)$$

158 where the integral is over the whole volume  $V$  of the climate system,  $\dot{q}_{rad}$  is the radiative  
 159 heating rate,  $\dot{s}_{mat}$  is the instantaneous specific rate of material entropy production due to  
 160 irreversible processes involving the climatic fluid, and  $T$  the temperature field. The following  
 161 expression is usually adopted for  $\dot{s}_{mat}$  [14, 16, 28, 41]:

$$\dot{s}_{mat} = \frac{\epsilon^2}{T} + \mathbf{F}_{SH} \cdot \nabla \left( \frac{1}{T} \right) + \mathbf{F}_{LH} \cdot \nabla \left( \frac{1}{T} \right) \quad (2)$$

162 where  $\epsilon^2$  the specific dissipation rate of kinetic energy,  $\mathbf{F}_{SH}$  the turbulent sensible heat  
 163 flux,  $\mathbf{F}_{LH}$  the turbulent latent heat flux, where by turbulent we mean not related to large  
 164 scale advection due to winds, which is in principle reversible. Romps [43] refers to the  
 165 representation of the entropy production given by Eq. (2) as resulting from the bulk heating  
 166 budget, because water is treated mainly as a passive substance, while processes such as  
 167 irreversible mixing of water vapor are altogether ignored. More detailed description of  
 168 the moist atmosphere have led to a consistent treatment of the entropy generated by the  
 169 various processes accounting for hydrological cycle these processes [8, 39, 40, 43]. Apparently,  
 170 though, the overall effect of hydrological cycle-related entropy production is captured quite  
 171 well using Eq. (2) [8, 30, 36].

172 Integrating the term  $\dot{s}_{mat}$  in Eq. (1) over the volume  $V$  of the climate system and taking  
 173 a long-term average, we obtain the average rate of material entropy production:

$$\overline{\dot{S}_{mat}} = \int_V d^3\mathbf{x} \overline{\dot{s}_{mat}} = \overline{\dot{S}_{mat}^{dir}}, \quad (3)$$

174 which gives the so called *direct formula* for the material entropy production. Using Eq. (1),  
 175 we derive an equivalent expression involving radiative heating rates only:

$$\overline{\dot{S}_{mat}} = - \int_V d^3\mathbf{x} \overline{\left(\frac{\dot{q}_{rad}}{T}\right)} = \overline{\dot{S}_{mat}^{ind}}, \quad (4)$$

176 which is the *indirect formula* for computing the average rate of entropy production, where  
 177 obviously  $\overline{\dot{S}_{mat}^{dir}} = \overline{\dot{S}_{mat}^{ind}} = \overline{\dot{S}_{mat}}$ . Equation (1) provides an intimate link between the radiative  
 178 fields and the material flow properties inside the climate system. Moreover, Eq. (4) is very  
 179 powerful because it permits to work out the average rate of material entropy production

180 by considering only the optical properties of the fluid. Pascale et al. [36] showed using the  
 181 climate model FAMOUS adopted in this study that  $\overline{\dot{S}_{mat}^{dir}}$  and  $\overline{\dot{S}_{mat}^{ind}}$  agree up to an excellent  
 182 degree of precision (within 1%). Since Pascale et al. [36] used the approximate expression  
 183 (2) for the specific material entropy production, they also confirmed that, indeed, at all  
 184 practical purposes using such simplified representation of the irreversibility associated to  
 185 the hydrological cycle is appropriate.

## 186 A. Entropy diagnostics in Climate Models

187 In an actual climate model the implementation of entropy diagnostics faces some difficul-  
 188 ties, both at theoretical level and in terms of practical implementation of the entropy-related  
 189 diagnostics. A theoretical difficulty is that, as evidenced in [33], many state-of-the-art cli-  
 190 mate models features an inconsistent energetics, such that when all parameters are held  
 191 fixed and the system reaches a steady state, the long-term average of the energy budget  
 192 at the top of the atmosphere (TOA), which is the only boundary of the climate system, is  
 193 unexpectedly biased with respect to the vanishing long-term average one should expect to  
 194 observe. Interestingly, all biased models feature a positive energy budget at TOA, which  
 195 implies that the time averaged outgoing long wave radiative flux is smaller than the net  
 196 incoming shortwave flux. This fact implies that there must be a positive definite spurious  
 197 sink of energy somewhere inside the system. More specific analyses make clear that such  
 198 spurious sinks are related to the imperfect closure of the hydrological cycle [24] and to the  
 199 inconsistent treatment of the dissipation of kinetic energy, which is not entirely (or at all)  
 200 fed back into the system as thermal energy [33]. Such inconsistencies at smallspatial and  
 201 temporal scales impact large scale, long term climatic properties. As a result, climate models  
 202 are biased cold, taking into consideration that the Earth emits approximately as black body,  
 203 or feature negative biases in the planetary albedo, or both. Moreover, since the biases are  
 204 related to climate processes, they are climate-dependent, and so hard to control a posteriori  
 205 via removal of anomalies. In terms of entropy production, an energy bias of the order of  
 206  $1 \text{ W m}^{-2}$  causes a bias in the entropy production of about  $4 \times 10^{-3} \text{ W m}^{-2} \text{ K}^{-1}$ , which  
 207 is comparable with the range of estimates of material entropy production given by various  
 208 climate models [30, 36]. The FAMOUS model we use in this study features minor inconsis-  
 209 tencies in terms of closure of the energy budget (the bias is smaller than  $0.1 \text{ W m}^{-2}$ , so that



210 the problem exposed here does not affect significantly our results (see discussion later).

211 Moreover, in a climate model it is hard to deal directly with Eq. (2) because material  
 212 turbulent fluxes are evaluated through parametrizations of unresolved processes. The cor-  
 213 responding routines in the numerical code do not give as outputs heat fluxes. On the other  
 214 hand the heating rates (i.e. the divergence of the heat fluxes) are easily diagnosed for these  
 215 unresolved processes. Neglecting the geothermal flux from the inner Earth and noting that  
 216 at the top-of-the-atmosphere we have only radiative fields, using Gauss' theorem, the mate-  
 217 rial entropy production can be worked out by considering all the material diabatic heating  
 218 rates, as shown in [36]:

$$\overline{\dot{S}_{mat}^{dir}} = \int_V d^3\mathbf{x} \left( \overline{\frac{\epsilon^2}{T}} \right) - \overline{\left( \frac{\nabla \cdot \mathbf{F}_{SH}}{T} \right)} - \overline{\left( \frac{\nabla \cdot \mathbf{F}_{LH}}{T} \right)} = \int_V d^3\mathbf{x} \left( \overline{\frac{\dot{q}_{mat}}{T}} \right) \quad (5)$$

219 In this paper we refer to the entropy budget of the FAMOUS GCM [15] which has been  
 220 studied in detail by [36, 37]. Lets first focus on the evaluation of  $\overline{\dot{S}_{mat}^{dir}}$ . Different pro-  
 221 cesses contribute to the entropy production terms described in Eq.(2): the heating rates  
 222 are calculated as output of many different parametrization routines describing the unre-  
 223 solved processes in the various subdomains of the climate system (atmosphere, ocean, soil,  
 224 cryosphere):

- 225 • Entropy production due to dissipation of kinetic energy,  $\overline{\dot{S}_{KE}}$ , defined as:

$$\overline{\dot{S}_{KE}} = \int d^3\mathbf{x} \left( \overline{\frac{\epsilon^2}{T}} \right). \quad (6)$$

226 In FAMOUS and, in general, in most climate models, the kinetic energy is dissipated  
 227 mainly through four parametrized processes: the turbulent stresses occurring at the  
 228 boundary layer, which extract kinetic energy from the free atmosphere, the gravity  
 229 wave drag, which dissipates kinetic energy in the upper atmosphere, atmospheric con-  
 230 vective processes, and small scale turbulence, which is represented by the horizontal  
 231 momentum hyperdiffusion (which serves also the purpose of increasing the numerical  
 232 stability of the model). In FAMOUS only the atmosphere contributes to this part of  
 233 the entropy production. This is a reasonable approximation because the dissipation  
 234 of kinetic energy occurring in the atmosphere is about two orders of magnitude larger  
 235 than that occurring in the ocean [41, 49].

- 236 • Entropy production due to irreversible transfer of sensible and latent heat via turbulent

237 fluxes,  $\overline{\dot{S}_{heat}} = \overline{\dot{S}_{SH}} + \overline{\dot{S}_{LH}}$ , defined as:

$$\overline{\dot{S}_{heat}} = \int d^3\mathbf{x} \left[ -\overline{\left(\frac{\nabla \cdot \mathbf{F}_{SH}}{T}\right)} - \overline{\left(\frac{\nabla \cdot \mathbf{F}_{LH}}{T}\right)} \right] = \overline{\dot{S}_{SH}} + \overline{\dot{S}_{LH}} \quad (7)$$

238 The boundary layer scheme contributes to the entropy production due to irreversible  
 239 sensible and latent heat transfer in the four subdomains of the climate system, as it  
 240 couples them through exchanges of sensible heat and water vapour; other parametrized  
 241 processes contributing to  $\overline{\dot{S}_{SH}}$  and  $\overline{\dot{S}_{LH}}$  are atmospheric convection and the conden-  
 242 sation and evaporation of water in the atmosphere, as determined by the clouds and  
 243 precipitation parametrization schemes. Instead, processes contributing only to  $\overline{\dot{S}_{SH}}$   
 244 are the oceanic convection, the small scale turbulent mixing of temperature described  
 245 by hyperdiffusion, and the mixing occurring inside the ocean associated to small scale  
 246 eddies and in the mixed layer.

247 Table I provides a synthetic outline of which routines describing unresolved processes con-  
 248 tribute to the various terms of the material entropy production in each climatic subdomain.  
 249 Therefore, in practice, we compute  $\overline{\dot{S}_{mat}^{dir}}$  as follows:

$$\overline{\dot{S}_{mat}^{dir}} = \sum_k \sum_c \int_{V_c} d^3\mathbf{x} \overline{\left(\frac{\dot{q}_k^c}{T}\right)} \quad (8)$$

250 where  $\dot{q}_k^c$  is the local instantaneous heating rate occurring in the subdomain  $V_c$  due to the  
 251 process  $k$ .

252 The evaluation of  $\overline{\dot{S}_{mat}^{ind}}$  is much easier because the heating rates are readily available from  
 253 the radiation scheme, which affects all the subdomains  $c$  of the climate system:

$$\overline{\dot{S}_{mat}^{ind}} = - \sum_c \int_{V_c} d^3\mathbf{x} \left[ \overline{\left(\frac{\dot{q}_{sw}^c}{T}\right)} + \overline{\left(\frac{\dot{q}_{lw}^c}{T}\right)} \right] \quad (9)$$

254 where we have divided the contribution  $\dot{q}_{sw}$  coming from the shortwave radiation, which  
 255 is only absorbed (and scattered), inside the climate systems, so that  $\dot{q}_{sw} \geq 0$ , from the  
 256 contribution  $\dot{q}_{lw}$  coming from the longwave radiation, which instead is scattered, absorbed,  
 257 and emitted, and is the sole responsible for the radiative cooling.

258 **III. COARSE-GRAINING OF THE ENTROPY PRODUCTION TERMS: DEFINITIONS AND SOME CONJECTURES**  
 259

260 The entropy budget is estimated using space and time integrals of the ratio between  
 261 the local heating term and the local temperature. In many cases, either because we need  
 262 to compress data or because climatological database only contain certain time or spatially  
 263 averaged data, we have to deal with coarse grained data for the heating rate  $\langle q(\mathbf{x}, t) \rangle_v^\tau$ , and  
 264 for the temperature  $\langle T(\mathbf{x}, t) \rangle_v^\tau$ , where  $\tau$  refers to the time scale of the temporal averaging  
 265 operation, and  $v$  refers to the set of stencil regions  $v(x)$  centered over  $x$  over which (mass-  
 266 weighted) spatial averaging is performed:

$$\langle X(\mathbf{x}, t) \rangle_v^\tau = \frac{1}{\tau \mu(v, t)} \int_v d^3 \mathbf{y} \int_{-\tau/2}^{\tau/2} d\sigma X(\mathbf{x} + \mathbf{y}, t + \sigma) \quad (10)$$

where  $\mu(v(x), t) = \int_{v(x)} d^3 \mathbf{x}$  is the mass contained in the stencil  $v(x)$  at time  $t$ . Mass-weighting is the natural choice in climate models as hydrostatic approximation is almost invariably used and vertical coordinates are expressed to a very good approximation in terms of pressure levels. Since the integrands in Eqs. (4) and (5) are nonlinear, we obviously have that for every  $\tau$  and  $v$ :

$$\overline{\dot{S}_{mat}^{dir}} = \int_V d^3 \mathbf{x} \overline{\left( \frac{\dot{q}_{mat}}{T} \right)} \neq \int_V d^3 \mathbf{x} \overline{\left( \frac{\langle \dot{q}_{mat} \rangle_v^\tau}{\langle T \rangle_v^\tau} \right)} = \overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}, \quad (11)$$

$$\overline{\dot{S}_{mat}^{ind}} = - \int_V d^3 \mathbf{x} \overline{\left( \frac{\dot{q}_{rad}}{T} \right)} \neq - \int_V d^3 \mathbf{x} \overline{\left( \frac{\langle \dot{q}_{rad} \rangle_v^\tau}{\langle T \rangle_v^\tau} \right)} = \overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}. \quad (12)$$

Moreover, while as discussed before  $\overline{\dot{S}_{mat}^{dir}} = \overline{\dot{S}_{mat}^{ind}}$ , there is no *a priori* reason to expect that  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  and  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  have the same value. Finally, we have that up to first order:

$$\overline{\dot{S}_{mat}^{dir}} - \overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau} = \Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_v^\tau \simeq - \int_V d^3 \mathbf{x} \frac{\Delta [\dot{q}_{mat}]_v^\tau \Delta [T]_v^\tau}{[\langle T \rangle_v^\tau]^2} \quad (13)$$

$$\overline{\dot{S}_{mat}^{ind}} - \overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau} = \Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_v^\tau \simeq \int_V d^3 \mathbf{x} \frac{\Delta [\dot{q}_{rad}]_v^\tau \Delta [T]_v^\tau}{[\langle T \rangle_v^\tau]^2} \quad (14)$$

267 where  $\Delta [X]_v^\tau = X - \langle X \rangle_v^\tau$ . It is natural to interpret  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$ ,  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  as the contribution  
 268 to the entropy production due to irreversible processes occurring on scales large than what  
 269 described by  $\tau$  and  $v$ . Consequently,  $\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_v^\tau$ ,  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_v^\tau$  in Eqs. (13)-(14) can be inter-  
 270 preted as the contributions to the entropy production given by the material flows (Eq. (13))  
 271 and radiative fluxes (Eq. (14)) with variability confined below the spatial scale given by  $v$   
 272 and by the time scale given by  $\tau$ .

273 Equations (11)-(14) address practical questions such as: what is the error related to  
 274 remapping the output of a climate model to a new resolution in space and time? How do  
 275 diurnal, seasonal and interannual variability and how different spatial structures (midlatitude  
 276 cyclones, equator-pole contrasts, longitudinal asymmetries due to ocean-land contrasts, etc)  
 277 affect the entropy budget? How should we proceed to compare the estimates of material  
 278 entropy production from models with different resolutions? Moreover, we need to understand  
 279 whether it is more accurate to obtain estimates of the material entropy production from  
 280 coarse grained fields of the radiative heating rates or of the material heating rates, which  
 281 can be used for the indirect or direct formula for the entropy production, respectively. These  
 282 issues may be sequentially investigated by filtering  $q(\mathbf{x}, t)$  and  $T(\mathbf{x}, t)$  over the associated  
 283 time- and space- scales.

284 When we perform the coarse graining given in Eq. (10) to the thermodynamic variables  
 285 and estimate the entropy production, we discount for the mixing processes occurring below  
 286 the chosen spatial and time scales. Therefore, one expects that  $\overline{\dot{S}_{mat}^{ind}/v}^\tau, \overline{\dot{S}_{mat}^{dir}}^\tau \geq 0$  and  
 287  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau, \Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau \geq 0$  for all choices of  $\tau$  and  $v$ . Moreover, it seems natural to conjecture  
 288 that if, given a model output, we choose a coarser graining, we should obtain a lower estimate  
 289 of the entropy production, because we neglect the impact of a larger set of irreversible  
 290 processes. In other terms, we should have that  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_{v_1}^{\tau_1} \geq \Delta \left[ \dot{S}_{mat}^{dir} \right]_{v_2}^{\tau_2}$  (or  $\overline{\dot{S}_{mat}^{dir}}_{v_1}^{\tau_1} \leq$   
 291  $\overline{\dot{S}_{mat}^{dir}}_{v_2}^{\tau_2}$ ) and  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_{v_1}^{\tau_1} \geq \Delta \left[ \dot{S}_{mat}^{ind} \right]_{v_2}^{\tau_2}$  (or  $\overline{\dot{S}_{mat}^{ind}}_{v_1}^{\tau_1} \leq \overline{\dot{S}_{mat}^{ind}}_{v_2}^{\tau_2}$ ) if  $\tau_2 \leq \tau_1$  and  $v_2 \subset v_1$ .  
 292 Let's see how to interpret the inequalities  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau, \Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau \geq 0$  using the r.h.s. of Eqs.  
 293 (13)-(14):

- 294 • The inequality  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau \geq 0$  can be interpreted as the fact that at all time and space  
 295 scales, there is on the global average a *positive* correlation between the anomalies of  
 296 radiative heating and the anomalies of temperature. This expresses the basic fact that  
 297 the climate system is driven by radiative forcings, in the first place. Hence, this term  
 298 refers to the *response of the system to the external forcing*. Note that the inequality  
 299 holds despite the the strong negative correlation between temperature anomalies and  
 300 long wave heating rate anomalies due to the Boltzmann feedback.
- 301 • The other inequality  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau \geq 0$ , instead, implies that at all time and space scales  
 302 on the average there is a *negative* correlation between the anomalies of heating due  
 303 to convergence of material heat fluxes and anomalies of temperatures. This relation,

304 instead, expresses the fact that temperature anomalies are damped by the geophysical  
 305 flows, and this terms refers to the *dissipation* occurring inside the system at all scales.

306 In other terms, these conjectured inequalities correspond to the well-known fact that the  
 307 climate is 1. forced by anomalies in the radiative forcing, and 2. the atmospheric and oceanic  
 308 circulations reduce the resulting temperature gradients. Climate processes are related in  
 309 such a way at all scales, *in the average*, i.e. when space and time averages are considered.  
 310 Obviously, locally in space and/or in time one can get, e.g. positive temperature fluctuations  
 311 and, at the same time, a positive heating due to latent heat release and sensible heat  
 312 convergence (e.g. tropical troposphere). Such processes can be positively correlated in time  
 313 for some locations, but this must come at the expenses of negative correlations dominating  
 314 elsewhere in the globe.

As the primary driving of climate is indeed the radiative forcing, while the fluid flows  
 tend to dampen the resulting temperature gradients through instabilities, Therefore, one  
 expects that the correlations between temperature and heating fields are stronger when  
 considering the radiative fields as sources of heating. In other terms, the convergence of  
 heat due to geophysical flows are neither strong nor fast enough to counter exactly the  
 radiative forcing at all scales. As an example, one may consider the fact that the radiative-  
 convective equilibrium is typically baroclinically unstable in the mid-latitudes, and, indeed,  
 baroclinic disturbances reduce the North-South temperature gradient by transporting heat  
 from South to North, but cannot reduce it to zero. Taking into consideration Eqs. (13)-(14),  
 the different role - forcings vs. dampening - of the convergence of the radiative fluxes vs.  
 material turbulent fluxes leads us to proposing an additional inequality. We conjecture that

$$\Delta \left[ \overline{S_{mat}^{ind}} \right]_v^\tau \geq \Delta \left[ \overline{S_{mat}^{dir}} \right]_v^\tau \quad \forall \tau, v,$$

from which, since  $\overline{S_{mat}^{dir}} = \overline{S_{mat}^{ind}}$ , we derive the following inequality

$$\overline{\langle S_{mat}^{dir} \rangle}_v^\tau \geq \overline{\langle S_{mat}^{ind} \rangle}_v^\tau, \quad \forall \tau, v.$$

#### 315 IV. RESULTS

316 We first discuss briefly how the coarse graining operation is performed in practice. Let  
 317 us consider a steady-state climate simulation lasting for a time period  $L$  (in our case  $L = 50$

318 years), which we divide it in  $N$  sub-intervals  $\tau = L/N$ , where  $\tau = M \times dt$ , where  $dt$  is the  
319 model's time step (1  $h$  in our case). The horizontal resolution is specified by regular grids  
320 with angle resolution of  $5^\circ \times 7.5^\circ$  lat-lon), while in the vertical we have 11 levels for the  
321 atmosphere, 20 oceanic levels, and 3 land surface levels [15]. Therefore, we subdivide the  
322 domain  $V$  of integration into  $Q$  subdomains  $v_q$ ,  $q = 1, \dots, Q$ , each containing (in the bulk of  
323 the model's domain)  $R$  grid points. Given an intensive thermodynamic field  $X(\mathbf{x}_k, t_j)$ , for  
324  $n = 1, \dots, N$  and  $q = 1, \dots, Q$ , we define its coarse grained version as:

$$\langle X(q, n) \rangle_v^\tau = \frac{1}{\tau \mu(v_q, n)} \sum_{j=R(q-1)+1}^{Rq} \sum_{k=M(n-1)+1}^{Mn} dt \mu(\mathbf{x}_j, \sigma_k) X(\mathbf{x}_j, \sigma_k), \quad (15)$$

325 where  $\mu(\mathbf{x}_j, \sigma_k) = \nu(\mathbf{x}_j) \rho(\mathbf{x}_j, \sigma_k)$  is the mass contained in the grid box centered around  $\mathbf{x}_j$  of  
326 volume  $\nu(\mathbf{x}_j)$  at time  $\sigma_k$  and  $\mu(v_q, n)$ , correspondingly, is the time averaged (for time ranging  
327 from  $t_{M(n-1)+1}$  and  $t_{Mn}$ ) mass contained in the domain  $v_q$  of volume  $\nu(v_q)$ . Therefore, our  
328 estimate of the coarse grained value of the material entropy production is:

$$\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau} = \frac{1}{N} \sum_{i=1}^Q \sum_{l=1}^N \nu(v_i) \frac{\langle \dot{q}_{mat}(i, l) \rangle_v^\tau}{\langle T(i, l) \rangle_v^\tau} \quad (16)$$

329 for the so-called direct formula, and:

$$\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau} = -\frac{1}{N} \sum_{i=1}^Q \sum_{l=1}^N \nu(v_i) \frac{\langle \dot{q}_{rad}(i, l) \rangle_v^\tau}{\langle T(i, l) \rangle_v^\tau} \quad (17)$$

330 for the so-called indirect formula. These formulas are the discretized versions of Eq. (11)  
331 and Eq. (12), respectively. Obviously, the discrete versions of the exact formulas for the  
332 material entropy production  $\overline{\dot{S}_{mat}^{dir}}$  and  $\overline{\dot{S}_{mat}^{ind}}$  are obtained by setting in Eq. (15)  $R = M = 1$ ,  
333 i.e., taking the model outputs at the highest possible resolution. The processes occurring  
334 in the interior of the ocean and below the first soil level, as these contributions have been  
335 shown to be entirely negligible in terms of entropy production [37], and so are discarded.

336 If we choose a given spatial resolution of our data and we consider different values of  
337  $M$ , we test how applying temporal coarse graining impacts the estimates of the material  
338 entropy production. Instead, if we change the shape of the stencil  $v$  and/or the number of  
339 points  $R$  while keeping  $M$  fixed, we investigate the impact of changing the spatial coarse  
340 graining scheme. Obviously, we cannot capture the contributions to the material entropy  
341 production due to irreversible processes taking place over timescale shorter than the model  
342 timestep and over space scales smaller than the model resolution. It is not clear, given a

343 specific model’s settings, how relevant these could be, and, indeed, the only way to find this  
 344 out is to alter the model’s resolution. This procedure may have relevance in terms of model  
 345 tuning, as one could decide to change a model’s parameter when altering its resolution in  
 346 such a way to keep the entropy production constant.

### 347 **A. Temporal coarse graining**

348 We start our investigation by performing coarse graining exclusively on time. We then  
 349 analyze a long, steady state model’s run lasting 50 years with a model’s timestep of 1 hour,  
 350 and consider 1 year as long-term averaging time. We use the following values for  $\tau$ : 1 hour  
 351 (model timestep,  $N = 1$ ), 6 hours ( $N = 6$ ), 12 hours ( $N = 12$ ), 1 day ( $N = 24$ ), 2 days  
 352 ( $N = 48$ ), 5 days ( $N = 120$ ), 10 days ( $N = 240$ ), 15 ( $N = 360$ ) days, 1 month ( $N = 720$ ),  
 353 3 months ( $N = 2160$ ), 6 months ( $N = 4320$ ), 1 year ( $N = 8640$ ). We then collect the  
 354 50 1-year averaged value of the coarse grained material entropy production and compute  
 355 the mean and standard deviation for the 50 data we have. Moreover, we consider longer  
 356 averaging periods - 5 years, 10 years, and 50 years, and take in these cases  $\tau$  equal to the  
 357 averaging time, so that  $N = 43200$ ,  $N = 86400$ , and  $N = 432000$  in the  $\tau = 5$ , 10, and  
 358 50 years case, respectively, thus spanning in total more than 5 orders of magnitude for  $N$ .  
 359 We then compute for the coarse-grained estimates of the material entropy production the  
 360 ten 5–year averages and the five 10–year averages, and compute the mean and standard  
 361 deviation, plus the unique value referred to the 50– year average. The statistics for such  
 362 large values of  $\tau$  are extremely stable.

363 In Fig.1(a) we report the estimates of the material entropy production obtained through  
 364 the direct formula  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_\tau}$  and the indirect formula  $\overline{\langle \dot{S}_{ind}^{dir} \rangle_\tau}$ , respectively, where we have  
 365 dropped the lower index  $v$  because we do not perform any spatial coarse graining. The  
 366 vertical bars indicate the uncertainty due to the long term variability.

367 The computed values (worked out at each timestep) of  $\overline{\dot{S}_{mat}^{ind}} \approx 53.1 \text{ mW m}^{-2} \text{ K}^{-1}$   
 368 ( $1 \text{ mW} = 10^{-3} \text{ W}$ ) and  $\overline{\dot{S}_{mat}^{dir}} \approx 53.5 \text{ mW m}^{-2} \text{ K}^{-1}$  have a difference of about 0.4 mW  
 369  $\text{m}^{-2} \text{ K}^{-1}$ , so that Eq. (1) is verified with great accuracy. The discrepancy term between the  
 370 two estimates is due to the extremely small spurious radiative imbalance at TOA of about  
 371  $0.1 \text{ W m}^{-2}$  (see [30]) and to numerical inaccuracies. Moreover, as discussed in [30, 37],  
 372 these estimates are in good agreement with what found in climate models of higher degree

373 of complexity.

374 As conjectured, we find that the estimates of the coarse grained entropy production  
375 decrease with increasing  $\tau$  from these reference values obtained with no temporal coarse  
376 graining. The bias resulting from the use of the indirect formula is larger for all values of  $\tau$ .  
377 In Fig. 1(a) we see that if we consider values of  $\tau$  up to 6 hours, the impact of coarse graining  
378 is extremely small. This implies that such small time scales the irreversible processes are  
379 negligible; this matches well with the fact that convection, which is the dominating fast  
380 process in the climate system, is parametrized with an instantaneous adjustment. This  
381 immediately points to an unwelcome spurious effects of climate parametrizations.

382 The effect of coarse graining becomes more relevant when  $\tau \geq 1$  day. Figures 1(b) and  
383 1(c) present the values of  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]^\tau$  and  $\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]^\tau$  as a function of  $\tau$ . For  $\tau \sim 1$  day, the  
384 difference between the true and the coarse grained value of the entropy production is about  
385  $\approx 0.6 \text{ mW m}^{-2} \text{ K}^{-1}$  if we use the direct formula, and  $\approx 2 \text{ mW m}^{-2} \text{ K}^{-1}$  is we use the  
386 indirect formula. Such biases are due to neglecting the mixing occurring on the time scale  
387 of the day, mostly due related to the daily cycle of incoming radiation. When considering  
388 the direct formula, it is interesting to note that  $\Delta \left[ \overline{\dot{S}_{KE}^{dir}} \right]^\tau$  is basically zero for all values of  
389  $\tau$  (not shown), meaning that there is no time correlation between the dissipation of kinetic  
390 energy and the temperature field. The coarse graining, instead, impacts the contribution to  
391 entropy production due to the hydrological cycle. We can substantiate this statement by  
392 observing that  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]^\tau \sim \Delta \left[ \overline{\dot{S}_{heat}^{dir}} \right]^\tau$  (see definition of the latter in Eq. 8), as can be seen  
393 by comparing Figs. 1(c) and 2(a).

394 The second timescale worth discussing is the one corresponding to 1 year ( $\sim 3 \times 10^7$  s).  
395 The use of annual means instead of time-step data introduces a bias of about  $4 \text{ mW m}^{-2} \text{ K}^{-1}$   
396 when using the indirect formula, which corresponds to neglecting the correlation between the  
397 seasonal cycle of the radiation budget and that of the radiation temperature field. Similarly,  
398 considering the direct formula, we obtain  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]^\tau \sim 1.5 \text{ mW m}^{-2} \text{ K}^{-1}$ , which measures  
399 the effect of neglecting the correlation of the seasonal cycle of the atmospheric and oceanic  
400 transport and dissipation and of the temperature field. One must note that a considerable  
401 contribution to the value of  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]^\tau$  for  $\tau \geq 1$  year is given by the atmospheric temperature  
402 hyperdiffusion, which in FAMOUS is implemented as a eight-order laplacian operator and  
403 applied after the advection to the model prognostic variables. Hyperdiffusion is generally  
404 introduced in dynamic cores for numerical reasons in order to smooth variables and avoid



405 local divergences. However it may thought as a way to represent turbulent dissipation and  
 406 mixing at subgrid scale. We discover that a traditional numerical *trick* used in the climate  
 407 modeling community for avoiding computational instabilities impacts a global scale physical  
 408 properties of the system, as observed in [33] when looking at energy budgets.

409 We also observe that there is no clear signature emerging in the functions  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]^\tau$  and  
 410  $\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]^\tau$  for values of  $\tau$  to weeks (synoptic waves) or monthly (low frequency variability)  
 411 time scales while a relatively smooth transitions is found going from daily to yearly averages.  
 412 This supports the idea that it is possible to look at weather disturbances as parts of a macro-  
 413 turbulent cascade.

414 Estimating the entropy production via either the direct or the indirect formula using  
 415 long term averages (but full spatial resolution) leads to underestimate the exact value of  
 416 the entropy production by less than 10%. This suggests that long term averages of the  
 417 climatic fields one can obtain from the climate repositories are enough to get a good idea  
 418 of the properties of the climate system. As we shall see in the next section, things change  
 419 drastically when the coarse graining impacts the spatial features of the climatic fields.

420 We conclude this section with a note on the oceanic processes, which we do not treat  
 421 in this paper as they contribute negligibly to the overall entropy production in the climate  
 422 system. in Fig. 2(a) we show the dependence of the entropy production due to the oceanic  
 423 mixing on the temporal coarse graining (dashed) line. We discover that its exact value,  
 424 computed at time step, is about  $1 \text{ mW m}^{-2} \text{ K}^{-1}$ , as in [37], and its coarse grained value  
 425 does not noticeably decrease up to  $\tau \sim 1$  year, above which the coarse grained estimate is  
 426 roughly halved. The dash-dotted line in Fig. 2(a) gives the contribution due to the vertical  
 427 mixing in the interior of the ocean, which is a very slow process and is, in fact, weakly  
 428 affected by the temporal coarse graining. The other contribution to the entropy production  
 429 in the ocean comes from the mixing occurring in the mixed layer. The mixing layer scheme  
 430 [21] parametrizes the convection due to heating at depth and cooling at the surface as well  
 431 as the mechanical stirring due to wind and is introduced in ocean models in order to account  
 432 for the seasonal thermocline variations. The coarse grained value of this term goes virtually  
 433 to zero for  $\tau \geq 1$  because, when considering such an averaging, we discount for the impact  
 434 of the seasonal cycle in the upper portion of the ocean.

## 435 **B. Spatial and Temporal coarse-graining**

436 In this section we analyze the combined effect of coarse graining the heating rates and  
437 temperature fields in space and time by using extensively Eqs. 16 and 17. Of course, there  
438 are many ways to perform coarse graining, boiling down to the selection of the stencil  $v$   
439 introduced before. Summarizing, we proceed as follows:

- 440 1. longitudinal averaging: the stencils  $v$  are given by arcs of varying length in the zonal  
441 direction;
- 442 2. areal averaging: the stencils  $v$  are given by portions of varying size of the spherical  
443 surface;
- 444 3. mass averaging: the stencils  $v$  are given by same-mass portions of the atmospheric  
445 spherical shell obtained by thickening in the vertical direction the stencils described  
446 in 2.;

447 in all cases we perform also temporal coarse graining by selecting the same averaging times  
448  $\tau$  described in the previous subsection.

449 It is important to note that the averaging as in points 1. and 2 is performed at constant  $x_3$ .  
450 FAMOUS (and HadCM3) uses hybrid vertical coordinates, *i.e.* a coordinate system which  
451 changes smoothly from a terrain-following specification near the lower boundary ( $\sigma$  coords.)  
452 to a isobaric definition ( $p$  coords.) in the medium-upper troposphere and stratosphere. As  
453 clear from Eqs. 15-17, the result of any coarse graining performed at constant value of  
454 the vertical coordinates depends on the vertical coordinate considered. In order to avoid  
455 the spurious effects of remapping the thermodynamic fields to a new coordinate system,  
456 we choose coarse grained grid boxes that respect as much as possible the original model's  
457 resolution.

458 We also remark that given the heavy computational burden of the operation, we restrict  
459 our analysis to only one of the fifty year of available data. We have tested that

### 460 *1. Longitudinal Averaging*

461 We first investigate the effect of coarse graining on the estimate of the material entropy  
462 production by averaging longitudinally the thermodynamic fields, up to the point of con-

463 sidering zonally averaged only fields, and by degrading their temporal resolution by using  
 464 the averaging times  $\tau$  described above. The estimates of  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  and  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  are given in  
 465 Fig. 3(a) and Fig. 3(c), respectively. The corresponding values of  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau$  and  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau$   
 466 are reported in Fig. 3(b) and Fig. 3(d), respectively, and some specific results are given in  
 467 Table II.

468 In all figures, the value of  $\tau$  is reported in the abscissae, while in the ordinates the value of  
 469 size of the spatial stencil, ranging from  $7.5^\circ$  (no coarse graining) to  $360^\circ$  (zonal averaging) is  
 470 shown. In both figures, the lower left corner corresponds to the best estimate of the entropy  
 471 production; the upper left corner corresponds to the material entropy production due to the  
 472 longitudinally averaged, high-temporal resolution fields. the lower right corner corresponds  
 473 to long-time averaged, high-resolution spatial case, and, eventually, the upper right corner  
 474 corresponds to the highest degree of coarse graining: it represent the entropy production due  
 475 to the long-term averaged, longitudinally averaged fields, and features the lowest value of  
 476  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  and  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$ . We remark that the values reported at the border of the domain given  
 477 by the lowest value of the ordinates coincide, obviously, with what shown in Fig. 1(a). As a  
 478 general fact, we observe that the coarse grained estimates of the entropy production decrease  
 479 (or remain virtually unchanged) as we perform coarser and coarser graining procedure, in  
 480 time or in space, and that  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau} \geq \overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$ .

481 The strongest dependence of the  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  is on  $\tau$ : temporal coarse graining appears to  
 482 be the dominating influence, while the effect of spatial coarse graining is apparent only  
 483 for  $\tau \leq 1$  day and for considerable longitudinal averaging, such that features below  $60^\circ$   
 484 are smeared out. In other terms, longitudinal averaging starts to matter only when we  
 485 lose information on the alternating pattern continents/oceans. The total effect of removing  
 486 totally the spatial structure is similar to that of performing a time-averaging of one day.  
 487 When considering  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  the picture is partially different: first, the influence of the spatial  
 488 averaging is relatively strong at all scales for  $\tau \leq 1$  day. The coupling between the spatial  
 489 and temporal scales indicates that using low pass filter and space and time we remove the  
 490 fast traveling synoptic waves of the mid-latitudes. As opposed to the case of the coarse  
 491 grained indirect estimate of the entropy production, spatial averaging plays a role also for 1  
 492 day  $\leq \tau \leq 3$  months. This is probably the signature of the relevance of low-frequency, large  
 493 scale features of the tropical circulation, which are sustained by longitudinal gradients (and  
 494 tend to reduce them), which are smeared out when extreme coarse graining is applied.

495 Concluding, we remark that neglecting information on the longitudinal fluctuations and  
 496 temporal fluctuations of the thermodynamic fields does not bias considerably (in the worst  
 497 case, by about 10%) the estimates of the entropy production one would obtain by retaining  
 498 the full information. This agrees with the fact that, in first approximation, in our planet  
 499 longitudinal gradients and longitudinal heat fluxes are relatively small [41]. Moreover, using  
 500 either the direct or indirect formula we obtain rather similar results, with a bias of maximum  
 501 5%. As expected, the direct formula gives more accurate estimates for all considered coarse  
 502 graining procedures.

## 503 2. Areal averaging

504 As a second step for understanding the role of spatial and temporal coarse graining of  
 505 the estimate of the material entropy production, we combine time averaging of the ther-  
 506 modynamic fields with areal averaging along horizontal surfaces. This operation allows us  
 507 to explore how the two-dimensional spatial covariance of heating and temperature fields  
 508 contributes to entropy production at all time scales. In order to keep coherence with the  
 509 previous coarse graining procedure, we proceed as follows. We divide the spherical surface in  
 510 coarse grained grid boxes defined by intervals (in degrees in latitude and longitude)  $(\Delta\lambda\Delta\phi)$   
 511 such that  $\Delta\phi/\Delta\lambda = 1.5$ , which is consistent with the model's resolution of  $5^\circ lat \times 7.5^\circ lon$ .  
 512 We then increase  $\Delta\phi$  from  $7.5^\circ$  up to  $90^\circ$ , thus decreasing progressively the number of coarse  
 513 grained grids from 1728 to 12. In order to complete the coarse graining, we select as two  
 514 coarsest resolutions  $(\Delta\lambda; \Delta\phi) = (90^\circ, 180^\circ)$  (four quadrants) and  $(\Delta\lambda; \Delta\phi) = (180^\circ, 360^\circ)$   
 515 (full spherical surface).

516 The estimates of  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  and  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  are given in Fig. 4(a) and Fig. 4(c), respectively. The  
 517 corresponding values of  $\Delta \left[ \overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau} \right]^\tau$  and  $\Delta \left[ \overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau} \right]^\tau$  are reported in Fig. 4(b) and Fig. 4(d),  
 518 respectively, and some specific results are given in Table II. We discover that, as opposed  
 519 to the previous case, the impact of selecting coarser and coarser graining in space reduces  
 520 considerably the value of  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  for all values of  $\tau$ , because such an averaging progressively  
 521 removes the strong meridional dependence of the thermodynamic fields, up to the extreme  
 522 case of  $v$  being the whole Earth's surface. In this case, the estimate of the entropy production  
 523  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau} \sim 47.2 \text{ mWm}^{-2}\text{K}^{-1}$ . Note that when considering very strong spatial averaging the  
 524 effect of changing  $\tau$  is negligible, because the spatial averaging alone reduces the temporal

525 correlations by mixing areas of the planet experiencing, e.g., different seasons. The  $\tau$ -  
 526 dependence of  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  is relevant only for  $\tau \leq 1$  day and spatial scales smaller than  $\sim$   
 527  $3 - 4 \times 10^6$  m, which is, like in the previous case, hints at the fact that when averaging  
 528 over large spatial and temporal scales, we remove the variability corresponding to synoptic  
 529 waves.

530 The function  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  has a qualitatively similar but quantitatively stronger dependence  
 531 on  $\Delta\phi$  and  $\tau$  with respect to  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$ : the bias in the estimate production is larger for all  
 532 the considered coarse graining. Additionally, the indirect formula is more strongly affected  
 533 by averaging over long time scales  $\tau$ , similar to what seen in Fig. 3(c), because the coupling  
 534 between the seasonal cycle of the radiative budget and the temperature fields is very strong.  
 535 Note that when we consider global or quasi-global spatial coarse graining, such effect disap-  
 536 pears, because averaging such large scales already removes a large part of the season cycle  
 537 signal. This is different from what reported in Fig. 3(c), because zonal averaging, obviously,  
 538 cannot remove the asymmetry between northern and southern hemisphere.

### 539 3. Mass Averaging

540 Finally, we perform spatial averaging along the horizontal and vertical direction and  
 541 for different values of  $\tau$  for the direct and indirect formula of entropy production. In this  
 542 way, we are able to ascertain the relevance of the processes involving irreversible fluxes  
 543 across temperature gradients along the vertical direction. Results are reported in Fig. 5  
 544 and Fig. 6, respectively, , and some specific results are given in Table III. Since we are  
 545 now dealing with three variables describing the coarse graining - the amplitude in latitude  
 546  $\Delta\phi$ , the number of levels  $n$ , and  $\tau$ , we present two cross sections obtained for  $\tau = 1$  day  
 547 (virtually indistinguishable from  $\tau = 1$  hour) and  $\tau = 1$  year, reported as panels (a) and (c),  
 548 respectively, in both Figs. 5 and 6. In panels (b) and (d) of these two figures, we report,  
 549 instead,  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau$  and  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau$ , respectively. The inequality  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau < \Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau$  is  
 550 clearly obeyed.

551 As we know from the previous discussions, the function  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v^\tau}$  is relatively weakly af-  
 552 fected by coarse graining along the horizontal directions and along the time axis. The modest  
 553 importance of time averaging is confirmed in this more complete analysis, as Figs. 5(a), and  
 554 5(c) are hard to distinguish. Instead, we find that averaging the thermodynamic fields along

555 the vertical reduces very severely the correlation between the temperature and heating fields,  
 556 so that the estimate of the entropy production obtained from the coarse grained fields is  
 557 much smaller than its true value  $\overline{\dot{S}_{mat}}$ . The results emphasize that vertical mixing is the  
 558 dominant effect contributing to the material entropy production.

559 When considering the indirect formula, we can draw roughly the same conclusions as  
 560 above, with the difference that  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  is more strongly affected by averaging along the  
 561 horizontal surface and along the time axis. Therefore, Figs. 6(a) and 6(c) feature clear  
 562 differences in terms of mean values, and, in each of them, the impact of performing very  
 563 coarse graining along the horizontal direction is more pronounced with respect to what  
 564 reported in Fig. 5. The effect of horizontal coarse graining become noticeable already when  
 565 grid boxes with side of  $(\Delta\lambda, \Delta\phi) \sim (20^\circ, 30^\circ)$  are considered. This implies that the spatial-  
 566 temporal correlation between radiative heating and temperature fields is relevant for larger  
 567 range of scales than in the case of described above. These results put in firmer ground the  
 568 results given in [30].

569 Comparing Figs. 5 and 6, one can verify that in all cases  $\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_v^\tau > \Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_v^\tau$ .  
 570 Moreover, one discovers that when considering the coarse possible graining, one obtains  
 571 that  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_{v_{M,v,h}}^{\tau_M}} \sim 16mWm^{-2}K^{-1} > \overline{\langle \dot{S}_{mat}^{ind} \rangle_{v_{M,v,h}}^{\tau_M}} \sim 0$ . In order to interpret some these  
 572 results, we shall consider some limiting cases for Eqs. (11)-(12). We first take as averaging  
 573 volume at each point at surface  $v = v_{M,v}$  the vertical column ranging from the bottom of  
 574 the fluid component of the climate system to the top of the atmosphere, and we consider a  
 575 long averaging time  $\tau = \tau_M \gg 1 y$ , so that all temporal dependencies are removed. In other  
 576 terms, we look at the thermodynamic properties of the *climatological fields*.

577 We start with the expression relevant for the indirect formula for estimating the material  
 578 entropy production:

$$\overline{\langle \dot{S}_{mat}^{ind} \rangle_{v_{M,v}}^{\tau_M}} = - \int_V d^3\mathbf{x} \overline{\left( \frac{\langle \dot{q}_{rad} \rangle_v^\tau}{\langle T \rangle_v^\tau} \right)} = - \int_\Sigma dx_1 dx_2 \frac{F_{TOA}(x_1, x_2)}{T_{di}(x_1, x_2)}. \quad (18)$$

579 where given the choice of  $\tau$ , the time averaging operation given by the overbar in Eq. 18  
 580 is immaterial. In Eq. 18  $F_{TOA}(x_1, x_2) = F_{TOA}^{SW}(x_1, x_2) - F_{TOA}^{LW}(x_1, x_2)$  is the climatological  
 581 average of the net radiative wave flux at the top of the atmosphere (positive when there is  
 582 net incoming radiation towards the planet), while the lower indices *LW* and *SW* indicate  
 583 the long wave and shortwave components, respectively.  $T_{di}(x_1, x_2)$  is the long term mean  
 584 of the vertical average of the fluid temperature. Such quantity can be closely approximated

585 by the emission temperature  $T_E(x_1, x_2) = (F_{TOA}^{LW}(x_1, x_2)/\sigma)^{1/4}$ , where  $\sigma$  is the Boltzmann's  
 586 constant. Note that since  $F_{TOA}$  and  $T_E$  are positively correlated (regions having a net posi-  
 587 tive incoming radiation are warmer), we have that  $F_{TOA}$  and  $1/T_E$  are negatively correlated.  
 588 Since  $\int_{\Sigma} dx_1 dx_2 F_{TOA} = 0$ , we derive that  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  in Eq. (18) is positive, as expected.

Equation (18) can also be given a different interpretation. We have that  $F_{TOA}(x_1, x_2)$  is equal to the divergence of enthalpy transport due to the large scale climatological atmospheric and oceanic flow, so that

$$\begin{aligned} F_{TOA}(x_1, x_2) &= \nabla_2 \cdot \int_{z_{surf}}^{TOA} dx_3 [\mathbf{J}_{lat}(\mathbf{x}) + \mathbf{J}_{dry}(\mathbf{x})] \\ &= \nabla_2 \cdot [\tilde{\mathbf{J}}_{lat}(x_1, x_2) + \tilde{\mathbf{J}}_{dry}(x_1, x_2)]. \end{aligned} \quad (19)$$

589 In the previous equation, we have indicated with  $\mathbf{J}_{lat}(\mathbf{x}) = L_w q(\mathbf{x}) \{v_1(\mathbf{x}), v_2(\mathbf{x})\}$  and  
 590  $\mathbf{J}_{dry}(\mathbf{x}) = [C_p T(\mathbf{x}) + g x_3] \{v_1(\mathbf{x}), v_2(\mathbf{x})\}$  the large scale, advective horizontal fluxes of la-  
 591 tent heat and of dry static energy, respectively, where  $L_w$  is the latent heat of evaporation  
 592 of water (taken as a constant for simplicity),  $q$  is the specific humidity, and  $C_p$  is the heat  
 593 capacity of air at constant pressure, and  $v_1$  and  $v_2$  indicate the two components of the hor-  
 594 izontal velocity field. Finally, the  $\tilde{\phantom{x}}$  sign refers to the vertically integrated fluxes. Inserting  
 595 the right hand side of Eq. (19) in Eq. (18), we conclude that Eq. (18) gives the entropy  
 596 produced by the large scale horizontal transport of the geophysical flows and approximates  
 597 the climate system as a purely 2D system featuring irreversible heat transport from warm  
 598 to cold regions. Looking in the bottom right corner of Fig. 6(c) (which corresponds to the  
 599 last entry in Table III), we obtain a value of  $\sim 6.5 \text{ mWm}^{-2}\text{K}^{-1}$  (note that the result is  
 600 virtually unaltered when averaging over any  $\tau \geq 1y$ ).

601 We can bring the previous example to a more extreme case. If the spatial stencil  $v$  is the  
 602 whole climate domain  $v_{M,h,v}$ , it is easy to derive that:

$$\overline{\langle \dot{S}_{mat}^{ind} \rangle_{v=V}^{\tau_M}} = \mu(v) \frac{\overline{\langle \dot{q}_{rad} \rangle_v^\tau}}{\overline{\langle T \rangle_v^\tau}} = -\mu(v) \frac{\int_{\Sigma} dx_1 dx_2 F_{TOA}(x_1, x_2)}{\overline{\langle T \rangle_v^\tau}} = 0, \quad (20)$$

603 because we have reduced the climate to a zero-dimensional system with a unique temperature  
 604 where absorbed and emitted radiation are equal. Such a system is at equilibrium, cannot do  
 605 any work, and cannot sustain any irreversible process. See Fig. 6(c) and penultimate entry  
 606 in Table III.

Let's now repeat the same coarse graining operations for the direct formula. We first consider  $v = v_{M,v}$ . In each location, when integrating vertically, the surface sensible heat fluxes

cancel out with the heating rates associated to sensible heat fluxes in the atmosphere, so that for this choice of coarse graining their contribution to the entropy production vanishes.

We obtain:

$$\begin{aligned} \overline{\langle \dot{S}_{mat}^{dir} \rangle_{v_{M,v}}^{\tau M}} &= \int_V d^3\mathbf{x} \overline{\left( \frac{\langle \dot{q}_{mat} \rangle_v^\tau}{\langle T \rangle_v^\tau} \right)} = \int_\Sigma dx_1 dx_2 \left[ \frac{L_w [P(x_1, x_2) - E(x_1, x_2)]}{T_{cli}(x_1, x_2)} + \frac{\tilde{\epsilon}^2(x_1, x_2)}{T_{cli}(x_1, x_2)} \right] \\ &= \int_\Sigma dx_1 dx_2 \left[ \frac{-\nabla_2 \cdot \tilde{J}_{lat}(x_1, x_2)}{T_{cli}(x_1, x_2)} + \frac{\tilde{\epsilon}^2(x_1, x_2)}{T_{cli}(x_1, x_2)} \right]. \end{aligned} \quad (21)$$

607 where  $L_w [P(x_1, x_2) - E(x_1, x_2)] = -\nabla_2 \cdot \tilde{J}_{lat}(x_1, x_2)$  as imposed by conservation of water  
608 mass, with  $P(x_1, x_2)$  and  $E(x_1, x_2)$  time-averaged values of precipitation and evaporation,  
609 respectively [41]. Furthermore, we indicate with  $\tilde{\epsilon}^2(x_1, x_2)$  the vertically integrated kinetic  
610 energy dissipation rate, and we choose, as a first approximation  $T_{cli}(x_1, x_2)$  as characteristic  
611 temperature defined as before. Therefore, Eq. (21) suggests that the bottom right corner  
612 of Fig. 5(c) corresponds to the sum of entropy produced by large scale transport of latent  
613 heat plus the entropy produced by the dissipation of kinetic energy. We find a value of  
614  $\sim 18 \text{ mWm}^{-2}\text{K}^{-1}$ . Considering that the entropy production due to large scale transport  
615 of sensible heat is much smaller than the corresponding contribution due to latent heat  
616 transport (from the precise calculation we get a factor of about 5 as ratio between the  
617 two terms) we can derive that the dissipation of kinetic energy contributes for about  $\sim 13$   
618  $\text{mWm}^{-2}\text{K}^{-1}$  to the total material entropy.

619 Interestingly, if we compute  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_{v_{M,h,v}}^{\tau M}}$ , we do not obtain a vanishing result. While the  
620 contribution to entropy production due to heat fluxes is eliminated, the contribution coming  
621 from the dissipation of kinetic energy is not removed by the operation of coarse graining:

$$\overline{\langle \dot{S}_{mat}^{dir} \rangle_{v=V}^\tau} = \mu(v) \frac{\overline{\langle \dot{q}_{mat} \rangle_v^\tau}}{\langle T \rangle_v^\tau} = \mu(v) \frac{\int_\Sigma dx_1 dx_2 \tilde{\epsilon}^2(x_1, x_2)}{\langle T \rangle_v^\tau} > 0. \quad (22)$$

622 Equation (22) gives, to a good degree of approximation, the *minimum* value of the entropy  
623 production compatible with the presence of a total dissipation  $\int_V d^3\mathbf{x} \overline{\epsilon^2(\mathbf{x}, t)}$  [28, 30]. The  
624 second entry of Table III reports for the contribution given in Eq. 22 a value of about 16  
625  $\text{mWm}^{-2}\text{K}^{-1}$ . This value agrees well with what derived using Eqs. 20-21 and with what  
626 obtained by direct estimate of  $\overline{\dot{S}_{KE}} \sim 13.5 \text{ mWm}^{-2}\text{K}^{-1}$ .

627 These results imply that we can obtain an extremely good estimate of the true value  
628 of the material entropy production even using climatological, horizontally global averages  
629 of the thermodynamics fields: in such a worst case scenario, we get a bias of about 10%.



630 Using few selected coarse grained estimates for the entropy production, one can derive that  
631 it is possible to split the total value of  $52.5 \text{ mWm}^{-2}\text{K}^{-1}$  as follows:  $\sim 6 \text{ mWm}^{-2}\text{K}^{-1}$   
632 can be attributed to large scale transports of latent and sensible heat;  $\sim 13 \text{ mWm}^{-2}\text{K}^{-1}$   
633 can be attributed to the entropy produced by dissipation of kinetic energy; the remaining  
634  $\sim 33.5 \text{ mWm}^{-2}\text{K}^{-1}$  can be attributed to vertical transports of sensible and latent heat  
635 (basically, convection). These estimates agree quite accurately with what obtained in [30]  
636 using scaling analysis. Also, one obtains, in agreement with the inequality proposed in [30],  
637 that the entropy produced by dissipation of kinetic energy is larger than that due to large  
638 scale energy transport.

## 639 V. CONCLUSIONS

640 The investigation of the climate system using tools borrowed from non-equilibrium ther-  
641 modynamics [2, 20] is a very active interdisciplinary research, which allows for connecting  
642 concepts of great relevance for climate dynamics, such as large scale heat transports and  
643 the Lorenz energy cycle [42], to basic thermodynamical concepts used in the investigation  
644 of general non-equilibrium systems [16, 28]. Such a theoretical framework seems relevant  
645 especially in the context of the growing field focusing on the study of the atmospheres of  
646 exoplanets [1, 32], for which detailed measurements are hardly available. In fact, ther-  
647 modynamical methods allow for defining inequalities and deducing apparently unexpected  
648 relations between different physical quantities [30].

649 In this paper we have focused on understanding where, in the Fourier space, dissipative  
650 and irreversible processes are dominant. This has been accomplished by computing how  
651 different spatial and temporal scales contribute to the material entropy production in the  
652 climate system. We have considered the output coming from a 50 *y* run under steady state  
653 conditions performed with the FAMOUS climate model [15], and have used the entropy  
654 diagnostics developed and tested in [36]. We have considered both the direct and the indirect  
655 formulas for material entropy production [8]: the former estimates the material entropy  
656 production using the heating rates associated to the dissipation of kinetic energy and the  
657 convergence of material heat fluxes, the latter uses, instead, the heating rates associated to  
658 radiative fluxes.

659 Our strategy has been the following: we have considered the estimates of the entropy

660 production coming from the coarse grained outputs of the heating and temperature fields.  
661 The coarse-graining has been performed both in time and in space. The temporal coarse  
662 graining ranges from hourly (timestep of the model) to yearly time scale, while the spatial  
663 coarse grained has been performed in three different modalities: 1) performing longitudinal  
664 averages; 2) performing averages along horizontal surfaces; 3) performing mass-weighted  
665 averages along the horizontal and vertical directions. We have conjectured and then verified  
666 numerically that the coarser the graining of the data, the lower is the resulting estimate of  
667 the material entropy production, both in the case of the direct and of the indirect formula  
668 for estimating the material entropy production. This implies that at all scales there is a  
669 negative correlation between heating rates related to flow (kinetic energy dissipation, sensible  
670 and latent heat fluxes) and the temperature field, and a positive correlation between the  
671 heating rate due to radiation and the temperature field. In other terms, at all scales, the  
672 climate systems results to be forced by radiation, while the resulting forced fluctuations are  
673 dissipated by the material fluxes. In agreement with this interpretation, we have conjectured  
674 that at all scales the correlation between the radiative heating and the temperature field is  
675 stronger than the correlation between the temperature field and the material heating rates.  
676 If the two correlation were equal, the climate system would be able to adjust, instantly and  
677 locally, to (spatial and temporal) variations in the radiative heating. The numerical results  
678 have provided support for this conjecture.

679 Considering various special cases of coarse graining, and using the basic thermodynamic  
680 equations, we have been able to estimate in a consistent way the contributions to material  
681 entropy productions coming from large scale horizontal heat transport ( $\sim 6 \text{ mWm}^2\text{-K}^{-1}$ ),  
682 dissipation of kinetic energy ( $\sim 13 \text{ mWm}^2\text{-K}^{-1}$ ), and vertical processes of sensible and  
683 latent heat exchanges (i.e. convection,  $\sim 33.5 \text{ mWm}^2\text{-K}^{-1}$ ). This suggest that, as first  
684 approximation, the climate system can be seen in terms of dissipative processes as a collection  
685 of weakly coupled vertical columns featuring turbulent exchanges and dissipation. This  
686 confirms the ideas presented in [30].

687 Note that one could use the quantitative information on the various contributions to the  
688 material entropy production to derive some basic properties of the climate system without  
689 resorting to the full three dimensional, time dependent fields, In particular, one can derive a  
690 good estimate of the intensity of the Lorenz energy cycle by multiplying the estimated value  
691 of the contribution of the dissipation to the kinetic energy to the material entropy production

692 times a characteristic temperature of the system, obtaining an estimate of  $\sim 3 \text{ Wm}^{-2}$ , which  
693 is in good agreement with what obtained after processing of the high-resolution data [36].

694 The fact that the estimate of the material entropy production of the climate system  
695 decreases when a coarser graining is considered is in qualitative agreement with what derived  
696 in the Appendix A for the simple case of continuum systems featuring generalized flux-  
697 gradient relations. This obviously does not imply that the climate system behaves as a  
698 diffusive system, yet share with a diffusive system this interesting property. The fact that at  
699 all spatial and temporal scales the system has a positive definite value of material entropy  
700 production, when global averages are considered. This is not in contradiction with the well-  
701 known phenomena of so-called *negative diffusion*, first noted by Starr [47], who observed  
702 that in certain portions of the atmosphere - namely, near the storm track - the fluxes of  
703 momentum transport momentum from low to high momentum regions. While this process  
704 - a crucial element of the general circulation of the atmosphere, observed in our model  
705 runs as well - seems to oppose the second law of thermodynamics, it is instead a local but  
706 macroscopic phenomenon, where the creation of organized structures (thanks to long-range  
707 correlations due to wave propagation), is, as we understand in this paper, over-compensated  
708 by large entropy production at the same scales elsewhere in the atmosphere.

709 Apart from providing insights on the properties of forced fluctuations and irreversible  
710 dissipative processes in the climate system at various spatial and temporal scales, this paper  
711 deals with the relationship between a model, its output, and the chosen observables, by  
712 providing information on what is the impact of being able to access data at lower resolution  
713 with respect to the model which has generated them. We have learnt that this lack of  
714 information always biases negatively our estimate of the entropy production, and that the  
715 bias is serious only if we miss information describing the vertical structure of thermodynamic  
716 fields.

717 Since performing time averages up to the yearly time scale does not bias substantially  
718 the estimates of the material entropy production, we have that it is possible to intercom-  
719 pare robustly the state-of-the-art climate models and assess on each of them the impact of  
720 climate change on the entropy production by resorting to the output data provided in the  
721 freely accessible PCMDI/CMIP3 ([http://www-pcmdi.llnl.gov/ipcc/about\\_ipcc.php](http://www-pcmdi.llnl.gov/ipcc/about_ipcc.php))  
722 and PCMDI/CMIP5 (<http://cmip-pcmdi.llnl.gov/cmip5/>) repositories, where long cli-  
723 mate runs outputs are typically stored in the form of monthly averaged data.

724 On a different note, the approach presented here seems promising specifically for the  
725 investigation of the atmosphere of exoplanets, because it allows for evaluating the error in  
726 the estimate of their thermodynamical properties due to the lack of high-resolution data.

727 In this paper, we have worked on the post-processing of data. The analysis presented  
728 here should be complemented with an additional investigation of how changing the reso-  
729 lution of a model impacts the estimate of its material entropy production, in total, and  
730 process by process. Using the material entropy production as cost function for addressing  
731 the interplay between the respective role of changes in the resolution of a model and of  
732 changes in the coarse graining of the post-processed data seems promising in the tantalizing  
733 quest for understanding what is a good model of a geophysical fluid and what is a robust  
734 parametrization. The results obtained here seem to have a much more general validity than  
735 for the specific case of the present Earth’s climate. Therefore, we plan to extend the present  
736 analysis, by studying the combined effect of changes in the resolution of the model and in  
737 the effective resolution of the post-processed in a simpler geophysical fluid dynamical system  
738 like an Aquaplanet. We believe that the conjectures presented on the effect of coarse grain-  
739 ing thermodynamic fields on the estimate of the material entropy production is of general  
740 validity for a vast range of systems that can be described by continuum mechanics.

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## Appendix A: Spectral Analysis of the Impact of Coarse Graining on the Entropy

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### Production for Diffusive Systems

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We wish to provide a simple outlook on how to interpret the results shown in this paper taking a spectral point of view. This is relevant in practical terms because many numerical models of geophysical fluids are actually implemented in spectral coordinates. We restrict ourselves to the contributions to the material entropy production coming from the presence of material fluxes transporting heat across temperature gradients. So, we do not consider here the term responsible for the dissipation of kinetic energy (which is weakly affected by coarse graining) nor the radiative terms contributing to the indirect formula. We consider the case of a much simpler simpler 3D continuum physical system, where heat transport obeys a generalized diffusive behavior. We make this choice not because we believe that the climate system is, in any real sense, diffusive, but because we wish to show that the observed dependence of the material entropy production on the coarse graining is in qualitative agreement with what would be obtained for a diffusive system.

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Let's assume that the contribution to the average rate of entropy production coming from the transport of heat due to the flow  $J$  across the temperature field  $T$  can be computed as:

$$\bar{S}_{mat}^J = \frac{1}{V} \frac{1}{T} \int_V d^3\mathbf{x} \int_0^T dt \vec{J} \cdot \vec{\nabla} \frac{1}{T} = -\frac{1}{V} \frac{1}{T} \int_V d^3\mathbf{x} \int_0^T dt \frac{\vec{J} \cdot \vec{\nabla} T}{T^2} = -\frac{1}{V} \frac{1}{T} \int_V d^3\mathbf{x} \int_0^T dt \frac{\vec{\nabla} \cdot \vec{J}}{T} \quad (\text{A1})$$

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Let's now make the simplifying diffusive-like assumption that  $\vec{J} = -\vec{\nabla} G(T)$ , where  $dG/dT > 0$ , so that the flux is always opposed in verse to the temperature gradient; in the usual linear flux-gradient approximation we have  $G(T) = \kappa T$ ,  $\kappa > 0$ . We derive:

$$\bar{S}_{mat}^J = \frac{1}{V} \frac{1}{T} \int_V d^3\mathbf{x} \int_0^T dt \frac{G'(T)}{T^2} |\vec{\nabla} T|^2 = \frac{1}{V} \frac{1}{T} \int_V d^3\mathbf{x} \int_0^T dt |\vec{\nabla} \Psi(T)|^2 > 0 \quad (\text{A2})$$

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where  $\Psi(T) = \int dT \sqrt{G'(T)/T^2}$ . For sake of simplicity - but without loss of generality - we assume that our domain  $\Sigma$  is a parallelepiped of sides  $L_x$ ,  $L_y$ , and  $L_z$ . Using Parseval's theorem, we derive that the rate of entropy production can be written as:

$$\bar{S}_{mat}^J = \sum_{p,q,r,s} (k_p^2 + k_q^2 + k_r^2) |\Psi_{p,q,r,s}|^2 = \sum_{p,q,r,s} S_{p,q,r,s}^{J,mat} \quad (\text{A3})$$

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where

$$\Psi_{p,q,r,s} = \frac{1}{V} \frac{1}{T} \int_V d^3\mathbf{x} \int_0^T dt \Psi \exp[2\pi i(p/L_x x + q/L_y y + r/L_z z - s/Tt)]. \quad (\text{A4})$$

772 Performing a spatio-temporal coarse graining to the field  $\Psi \rightarrow \tilde{\Psi}$  can be reframed as applying  
773 a linear filter in Fourier space, as  $\Psi_{p,q,r,s} \rightarrow \tilde{\Psi}_{p,q,r,s} = \Phi_{p,q,r,s} \Psi_{p,q,r,s}$ , where  $|\Phi_{p,q,r,s}| \leq 1$   
774  $\forall p, q, r, s$  plus the usual complex conjugacy properties. Note that in our case we would like  
775 to be able to apply the filtering to the  $T$  field and to the heating field  $-\vec{\nabla} \cdot \vec{J}$  and not to the  
776 function  $\Psi$  constructed here. Nonetheless, assuming that the relative change of  $T$  across the  
777 domain (see also Eq. (13)) is small, the conclusions are virtually unaltered.

778 Equation A3 has the remarkable property that all of terms of the summation  $S_{p,q,r,s}^{J,mat}$  are  
779 positive. We can interpret  $S_{p,q,r,s}^{J,mat}$  as the entropy produced by processes occurring at the  
780 scales described by the indices, in this case  $\lambda_x = L_x/p$ ,  $\lambda_y = L_y/q$ ,  $\lambda_z = L_z/r$ , and  $\tau = T/s$ .  
781 Therefore, indicating as  $\tilde{S}_{mat}^J$  the value of the entropy production for the coarse grained  
782 fields, we obtain:

$$\tilde{S}_{mat}^J = \sum_{p,q,r,s} (k_p^2 + k_q^2 + k_r^2) |\Phi_{p,q,r,s}|^2 |\Psi_{p,q,r,s}|^2 = \sum_{p,q,r,s} |\Phi_{p,q,r,s}|^2 S_{p,q,r,s}^{J,mat} \leq \bar{S}_{mat}^J. \quad (\text{A5})$$

783 In particular, we can associate a coarse graining on the scales  $\Lambda_x$ ,  $\Lambda_y$ ,  $\Lambda_z$ , and  $\tau$  (referred to  
784 the  $x$ -,  $y$ -,  $z$ -directions and time, respectively) to a filter of the form  $\Phi_{p,q,r,s} = 0$  if  $p > L_x/\Lambda_x$ ,  
785 or  $q > L_y/\Lambda_y$ , or  $r > L_z/\Lambda_z$ , or  $s > T/\tau$  and  $\Phi_{p,q,r,s} = 1$  otherwise. Slightly different ways  
786 of doing the coarse graining will result into different filters, which will be, nonetheless,  
787 asymptotically equivalent if the involved scales are the same.

788 The main conceptual point behind this result is independent of the shape of the of the  
789 domain of integration: the natural orthogonal expansion for atmospheric fields defined in  
790 an (approximately) spherical thin shell is given by spherical harmonics in for the latitudinal  
791 and longitudinal dependence and the usual Fourier expansion for the vertical direction. In  
792 the case of a thin spherical shell of thickness  $L_z$  situated at distance  $R$  from the center of  
793 the sphere, Eq. A3 can be rewritten as:

$$\bar{S}_{mat}^J = \sum_n \sum_{l \geq 0} \sum_{m=-l}^l \sum_s (k_n^2 + l(l+1)/R^2) |\Psi_{n,l,m,s}|^2 \quad (\text{A6})$$

794 with

$$\Psi_{n,l,m,s} = \frac{1}{L_z} \frac{1}{4\pi} \frac{1}{T} \int_{\Omega} d\Omega \int_0^T \Psi(z, \theta, \phi, t) \exp[2\pi i(n/L_z z - s/Tt)] * Y(\theta, \phi)_n^{m*}. \quad (\text{A7})$$

795 where  $Y(\theta, \phi)_n^m$  are the usual spherical harmonics and  $\Omega$  refers to the solid angle. In this  
796 case, performing a spatio-temporal coarse graining to the field  $\Psi \rightarrow \tilde{\Psi}$  results in reduced

797 value of the estimate of the entropy production:

$$\tilde{S}_{mat}^J = \sum_n \sum_{l \geq 0} \sum_{m=-l}^l \sum_s (k_n^2 + l(l+1)/R^2) |\Phi_{n,l,m,s}|^2 |\Psi_{n,l,m,s}|^2 \leq \bar{S}_{mat}^J. \quad (\text{A8})$$

798 It is now easy to relate the expression of  $|\Phi_{n,l,m,s}|^2$  to common averaging operation performed  
 799 on climate data. A coarse graining on the vertical scale  $\Lambda_z$ , on a temporal scale  $\tau$  and on a  
 800 horizontal surface area  $\sigma$  (or angular resolution  $\sigma/R^2$ ) amounts to setting  $|\Phi_{n,l,m,s}|^2 = 0$  if  
 801  $l \geq \sqrt{8\pi R^2/\sigma}$  (corresponding approximately to a triangular truncation  $T(K)$ , where  $K$  is  
 802 the integer closest to  $\sqrt{8\pi R^2/\sigma}$ ), or  $n > L_z/\Lambda_z$ , or  $s > T/\tau$ , and  $|\Phi_{n,l,m,s}|^2 = 1$  otherwise.  
 803 Instead, performing zonal averages corresponds to setting,  $|\Phi_{n,l,m,s}|^2 = 0$  if  $m \neq 0$ .

804 The bottom line of the previous considerations is that adopting a coarser graining cor-  
 805 responds to increasing the involved scales determining the spectral cutoff. if we assume a  
 806 flux-gradient relationship which is consistent with the second law of thermodynamics (even  
 807 if it is not the usual Fickian, linear relation), Eqs. A5 and A8 imply that as the graining  
 808 becomes coarser, the estimate of the entropy production becomes smaller, because we the  
 809 summation is performed over fewer terms, all of them positive. This behavior is independent  
 810 of the physical domain under consideration. Moreover, the previous considerations qualita-  
 811 tively apply - even if results are somewhat more cumbersome - if the relationship between flux  
 812 and gradient is more general than what previously assumed, e.g. if  $J_i = -\partial_i G_i(T)$  (where  
 813 the Einstein summation convention is not taken), under the condition that  $dG_i/dT < 0 \forall i$ .

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TABLE I. List of the main FAMOUS' parametrization routines for unresolved processes and their impact in term of heating rates  $\dot{q}_k^c$  on the various terms contributing to  $\dot{s}_{mat}$ . Codes: BL=Boundary Layer; AC=Atmospheric Convection; HD=Hyperdiffusion; OC=Oceanic Convection; D=Diffusion; GW=Gravity Waves; C/E=Condensation/Evaporation; ML=Mixed Layer

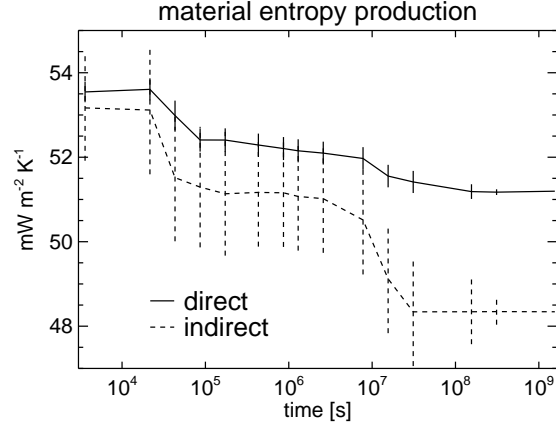
Contribution to $\dot{s}_{mat}$	Atmosphere	Ocean	Soil	Cryosphere
$\frac{\epsilon^2}{T}$	BL, C, GW, HD			
$-\frac{\nabla \cdot \mathbf{F}_{LH}}{T}$	BL, AC, C/E,	BL	BL	BL
$-\frac{\nabla \cdot \mathbf{F}_{SH}}{T}$	BL, AC, HD	BL, ML, OC, D	BL	BL

TABLE II. Values of  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_v^\tau$  and  $\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_v^\tau$  obtained when considering the coarsest resolution in space (lower index:  $v_M$ ), in time (lower index:  $\tau_M$ ), or both. Only 2D horizontal averagings are considered here. Reference values for highest resolution data:  $\overline{\dot{S}_{mat}^{dir}} = 52.5 \text{ mWm}^{-2}\text{K}^{-1}$  and  $\overline{\dot{S}_{mat}^{ind}} = 52.1 \text{ mWm}^{-2}\text{K}^{-1}$ . All values are in units of  $\text{mWm}^{-2}\text{K}^{-1}$ .

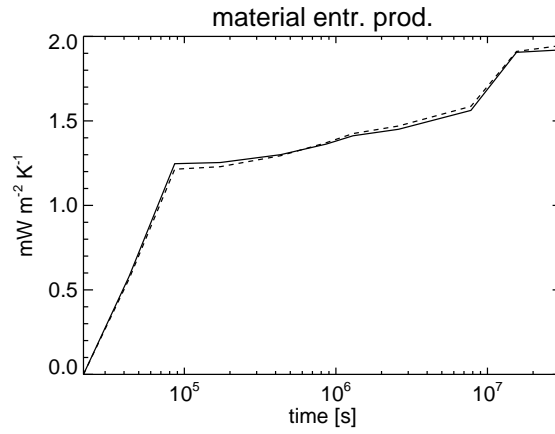
Averaging	$\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_{v_M}^\tau$	$\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_{v_M}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_{v_M}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_{v_M}^\tau$	$\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_{v_M}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_{v_M}^{\tau_M}$
Longitudinal	2.1	2.2	2.1	2.2	5.0	5.0
Surface	5.3	5.3	2.2	12.3	12.8	5.0

TABLE III. Values of  $\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_v^\tau$  and  $\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_v^\tau$  obtained when considering the coarsest resolution either in horizontal direction (lower index:  $v_{M,h}$  or vertical direction (lower index:  $v_{M,v}$ ), or in both (lower index:  $v_{M,h,v}$ ).  $\tau$  is set to 1 y. Reference values for highest resolution data:  $\overline{\dot{S}_{mat}^{dir}} = 52.5 \text{ mWm}^{-2}\text{K}^{-1}$  and  $\overline{\dot{S}_{mat}^{ind}} = 52.1 \text{ mWm}^{-2}\text{K}^{-1}$ . All values are in units of  $\text{mWm}^{-2}\text{K}^{-1}$ .

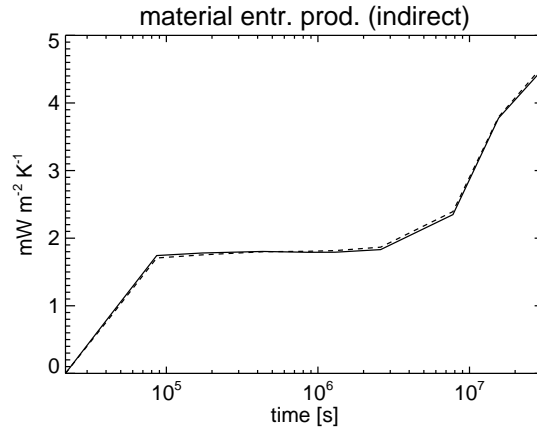
Averaging	$\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_{v_{M,h}}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_{v_{M,h,v}}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{dir}} \right]_{v_{M,v}}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_{v_{M,h}}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_{v_{M,h,v}}^{\tau_M}$	$\Delta \left[ \overline{\dot{S}_{mat}^{ind}} \right]_{v_{M,v}}^{\tau_M}$
Mass weighted	5.3	36.5	34.5	12.8	52.1	45.6



(a)

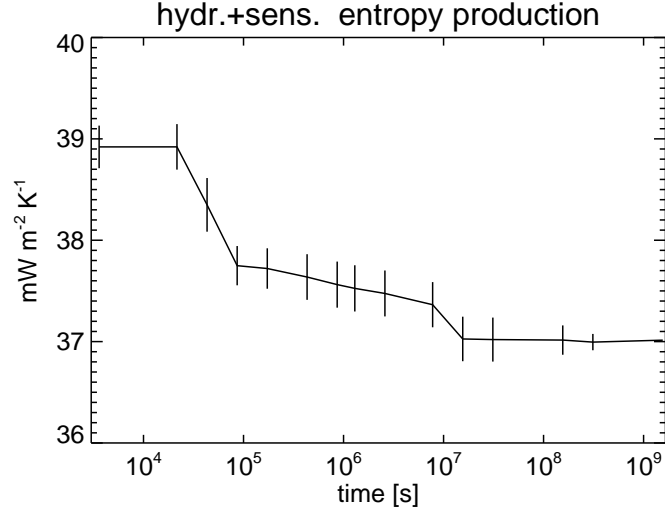


(b)

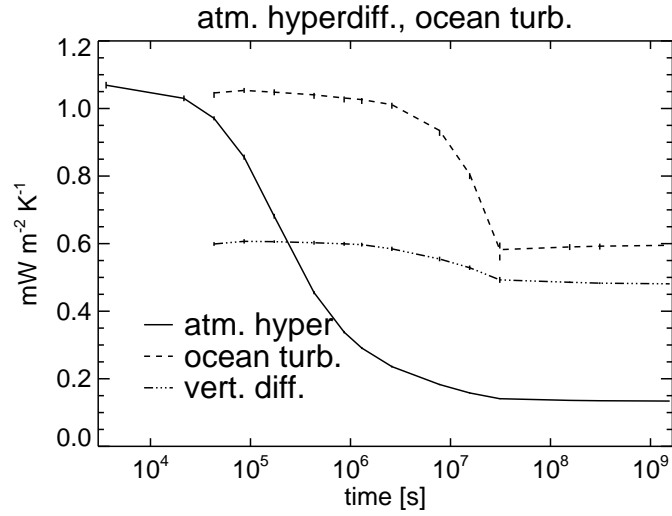


(c)

FIG. 1. (a) Time-scale dependence of the total material entropy production (direct and indirect estimates);  $\tau$  ranges between 3600 s (model timestep) to  $1.5 \times 10^9$  s (50 years); (b) Differences between the exact and time-averaged material entropy production  $\Delta \left[ \dot{S}_{mat}^{dir} \right]^\tau$  for  $3600 \text{ s} \leq \tau \leq 1 \text{ year}$  (c); as in (b) but for  $\Delta \left[ \dot{S}_{mat}^{ind} \right]^\tau$ . The dashed lines represent the correlation terms given in Eqs (14)-(15).



(a)



(b)

FIG. 2. (a) Material entropy production due to hydrological cycle and heat diffusion. Here  $L = 50$  years. (b) Material entropy production due to atmospheric small-scale temperature diffusion (continuous line). The material entropy production due to ocean turbulence (vertical and horizontal diffusion, mixed layer physics and convection is also reported, see [36] for details). We also show the vertical diffusion contribution (dotted-dashed line) to the total turbulent material entropy production. Note that the oceanic processes have a 12-hour timestep and have not been considered in the total entropy budget .

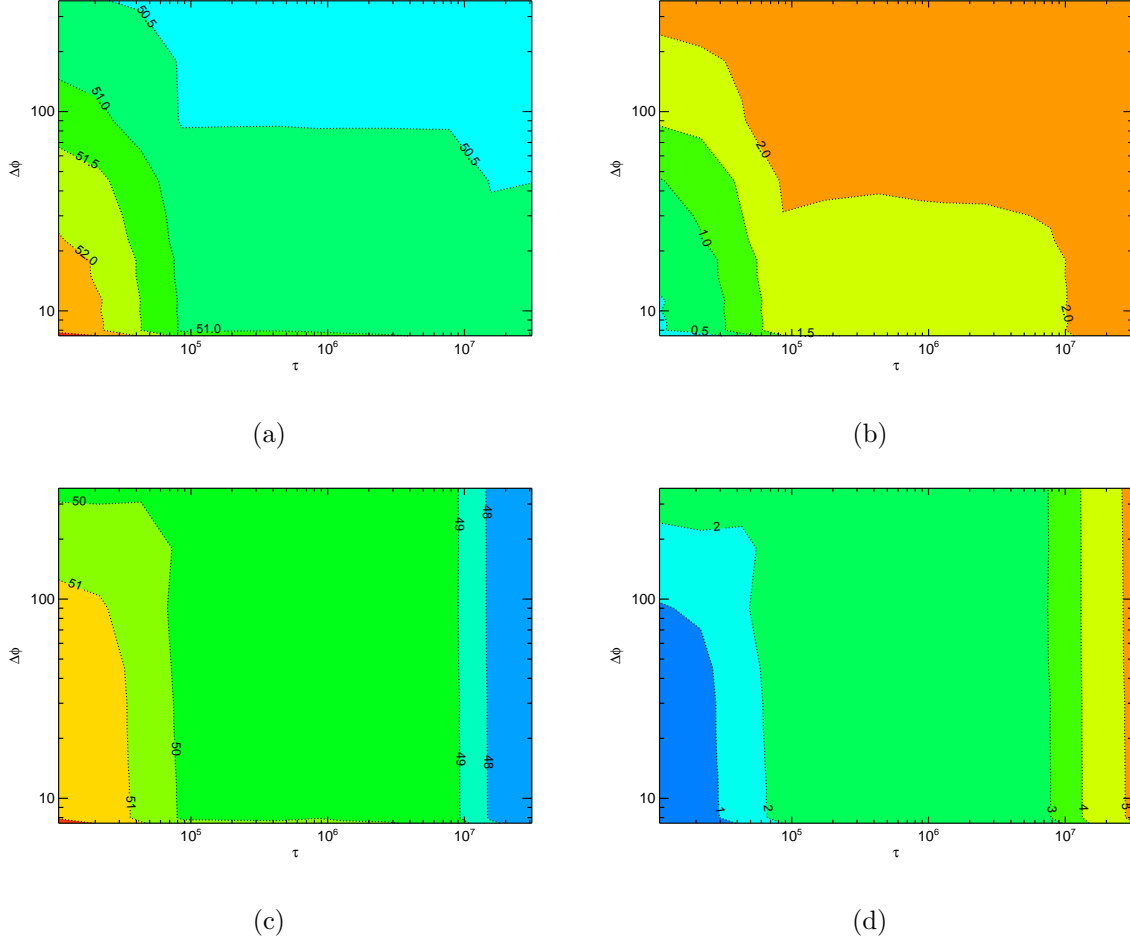
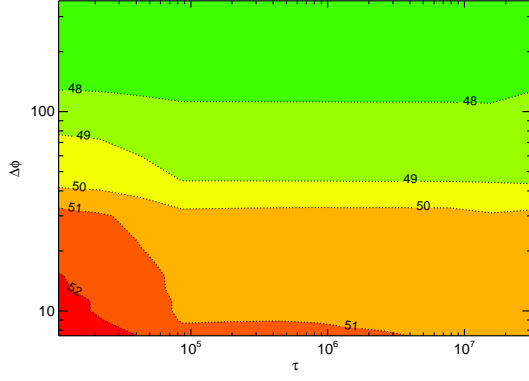
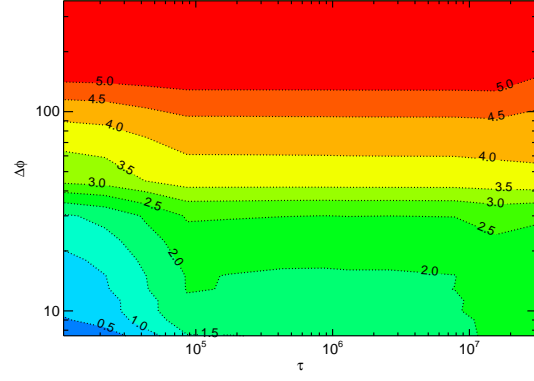


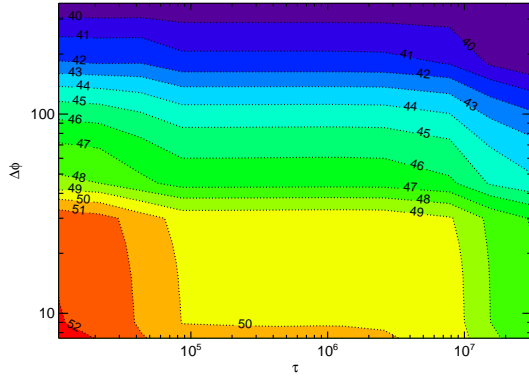
FIG. 3. Estimates of  $\langle \dot{S}_{mat}^{dir} \rangle_v^\tau$  (a),  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau$  (b),  $\langle \dot{S}_{mat}^{ind} \rangle_v^\tau$  (c), and  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau$  (d). The spatial averaging considered here is given by longitudinal averages on horizontal surface. The  $x$ -axis reports  $\tau$ , the  $y$ -axis describes the extent  $\Delta\phi$  of the averaging in  $^\circ$ .



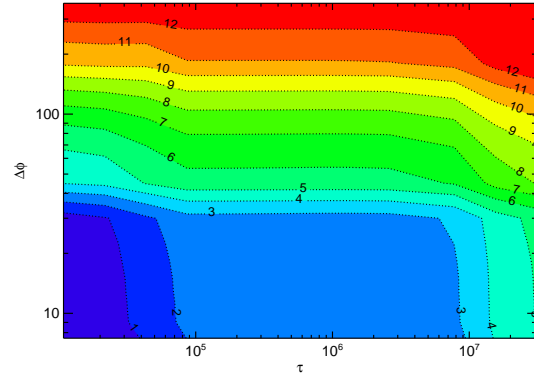
(a)



(b)



(c)



(d)

FIG. 4. Estimates of  $\overline{\dot{S}_{mat}^{dir}}_{\tau}$  (a),  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_{\tau}$  (b),  $\overline{\dot{S}_{mat}^{ind}}_{\tau}$  (c), and  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_{\tau}$  (d). The spatial averaging considered here is given by areal averages along horizontal surfaces. The  $x$ -axis reports  $\tau$ , the  $y$ -axis describes the longitudinal extent  $\Delta\phi$  of the coarse grained grid boxes. Details are given in the text.



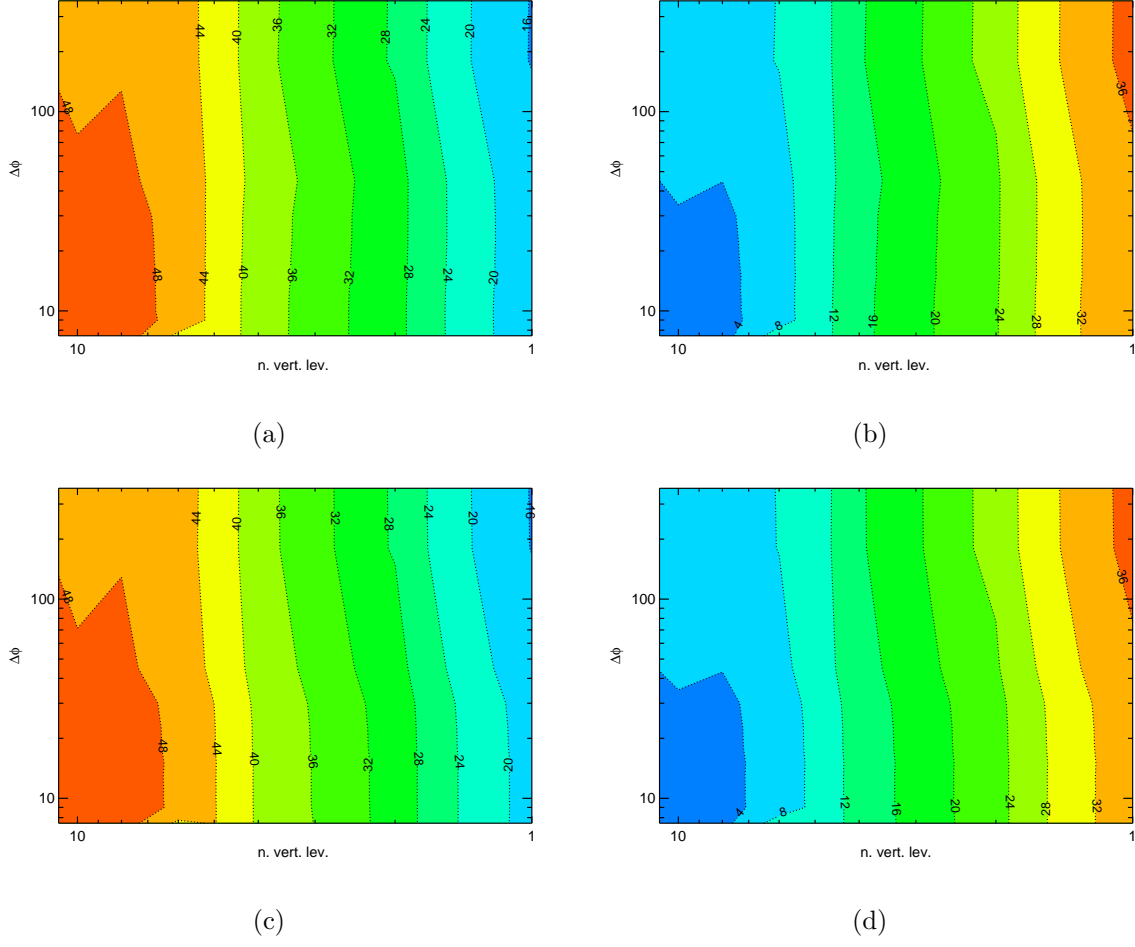
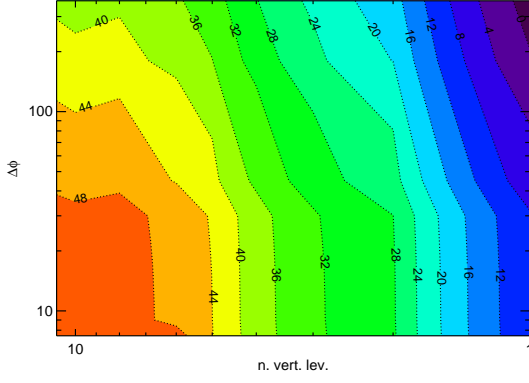
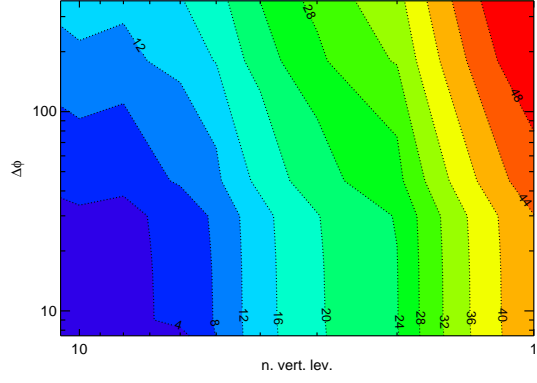


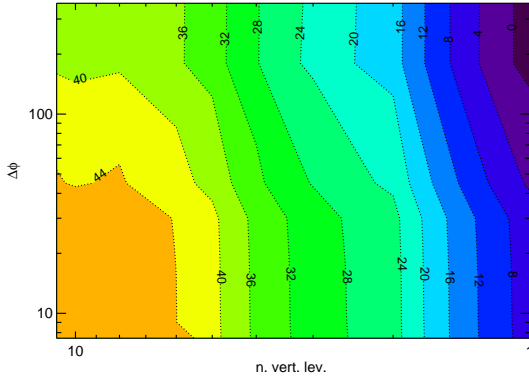
FIG. 5. Estimates of  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v}^\tau$  (a), and  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau$  (b) for  $\tau = 1$  day, and of  $\overline{\langle \dot{S}_{mat}^{dir} \rangle_v}^\tau$  (c), and  $\Delta \left[ \dot{S}_{mat}^{dir} \right]_v^\tau$  (d) for  $\tau = 1$  year. The spatial averaging considered here is given by areal averages along horizontal surfaces and vertical averages along columns. The  $x$ -axis reports the numbers of vertical levels involved in the averaging, the  $y$ -axis describes the longitudinal extent  $\Delta\phi$  of the coarse grained grid boxes. Details are given in the text.



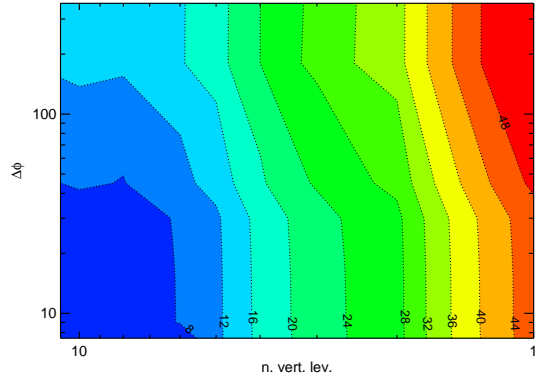
(a)



(b)



(c)



(d)

FIG. 6. Estimates of  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  (a), and  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau$  (b) for  $\tau = 1$  day, and of  $\overline{\langle \dot{S}_{mat}^{ind} \rangle_v^\tau}$  (c), and  $\Delta \left[ \dot{S}_{mat}^{ind} \right]_v^\tau$  (d) for  $\tau = 1$  year. The spatial averaging considered here is given by areal averages along horizontal surfaces and vertical averages along columns. The  $x$ -axis reports the numbers of vertical levels involved in the averaging, the  $y$ -axis describes the longitudinal extent  $\Delta\phi$  of the coarse grained grid boxes. Details are given in the text.